# Fermionic superstring loop amplitudes in the pure spinor formalism 

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Abstract: The pure spinor formulation of the ten-dimensional superstring leads to manifestly supersymmetric loop amplitudes, expressed as integrals in pure spinor superspace. This paper explores different methods to evaluate these integrals and then uses them to calculate the kinematic factors of the one-loop and two-loop massless four-point amplitudes involving two and four Ramond states.

Keywords: Superstrings and Heterotic Strings, Superspaces.

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## 1. Introduction

The quantisation of the ten-dimensional superstring using pure spinors as world-sheet ghosts []] has overcome many difficulties encountered in the Green-Schwarz (GS) and Ramond-Neveu-Schwarz (RNS) formalisms. Most notably, by maintaining manifest spacetime supersymmetry, the pure spinor formalism has yielded super-Poincaré covariant multiloop amplitudes, leading to new insights into perturbative finiteness of superstring theory (2), 3 .

Counting fermionic zero modes is a powerful technique in the computation of loop amplitudes in the pure spinor formalism and has for example been used to show that at least four external states are needed for a non-vanishing massless loop amplitude (2]. Furthermore, the structure of massless four-point amplitudes is relatively simple because all
fermionic worldsheet variables contribute only through their zero modes. In the expressions derived for the one-loop [2] and two-loop [4] amplitudes, supersymmetry was kept manifest by expressing the kinematic factors as integrals over pure spinor superspace [5] involving three pure spinors $\lambda$ and five fermionic superspace coordinates $\theta$,

$$
\begin{align*}
& K_{1 \text {-loop }}=\left\langle(\lambda A)\left(\lambda \gamma^{m} W\right)\left(\lambda \gamma^{n} W\right) \mathcal{F}_{m n}\right\rangle, \\
& K_{2 \text {-loop }}=\left\langle\left(\lambda \gamma^{m n p q r} \lambda\right)\left(\lambda \gamma^{s} W\right) \mathcal{F}_{m n} \mathcal{F}_{p q} \mathcal{F}_{r s}\right\rangle, \tag{1.1}
\end{align*}
$$

where the pure spinor superspace integration is denoted by $\langle\ldots\rangle$, and $A_{\alpha}(x, \theta), W^{\alpha}(x, \theta)$ and $\mathcal{F}_{m n}(x, \theta)$ are the superfields of ten-dimensional Yang-Mills theory.

The kinematic factors in (1.1) have been explicitly evaluated for Neveu-Schwarz states at two loops [6] and one loop [7], and were found to match the amplitudes derived in the RNS formalism [8]. This provided important consistency checks in establishing the validity of the pure spinor amplitude prescriptions. (Related one-loop calculations had been reported in [包].)

In this paper, it will be shown how to compute the kinematic factors in (1.1) when the superfields are allowed to contribute fermionic fields, as is relevant for the scattering of fermionic closed string states as well as Ramond/Ramond bosons. It turns out that the calculation of fermionic amplitudes presents no additional difficulties, making (1.1) a good practical starting point for the computation of four-point loop amplitudes in a unified fashion. This practical aspect of the supersymmetric pure spinor amplitudes was also emphasised in [10], where the tree-level amplitudes were used to construct the fermion and Ramond/Ramond form contributions to the four-point effective action of the type II theories.

This paper is organised as follows. In section 2, different methods to compute pure spinor superspace integrals are explored. These methods are then applied to the explicit evaluation of the kinematic factors of massless four-point amplitudes at the one-loop level in section 3, and at the two-loop level in section 4. In both these sections, the bosonic calculations are briefly reviewed before separately considering the cases of two and four Ramond states. Particular attention will be paid to the constraints imposed by simple exchange symmetries. An appendix contains algorithms which were used to reduce intermediate expressions encountered in the amplitude calculations to a canonical form.

## 2. Zero mode integration

The calculation of scattering amplitudes in the pure spinor formalism leads to integrals over zero modes of the fermionic worldsheet variables $\lambda$ and $\theta$. Both $\theta$ and $\lambda$ are 16-component Weyl spinors, the $\lambda$ are commuting and the $\theta$ anticommuting, and $\lambda$ is subject to the pure spinor constraint $\left(\lambda \gamma^{m} \lambda\right)=0$. The amplitude prescriptions [1], [2] require three zero modes of $\lambda$ and five zero modes of $\theta$ to be present, and a Lorentz covariant object

$$
\begin{equation*}
\overline{\mathcal{T}}^{\alpha \beta \gamma, \delta_{1} \ldots \delta_{5}} \equiv\left\langle\lambda^{\alpha} \lambda^{\beta} \lambda^{\gamma} \theta^{\delta_{1}} \ldots \theta^{\delta_{5}}\right\rangle=\overline{\mathcal{T}}^{(\alpha \beta \gamma),\left[\delta_{1} \ldots \delta_{5}\right]} \tag{2.1}
\end{equation*}
$$

was constructed such that the Yang-Mills antighost vertex operator

$$
\begin{equation*}
V=\left(\lambda \gamma^{m} \theta\right)\left(\lambda \gamma^{n} \theta\right)\left(\lambda \gamma^{p} \theta\right)\left(\theta \gamma_{m n p} \theta\right) \quad \text { has } \quad\langle V\rangle=1 \tag{2.2}
\end{equation*}
$$

In this section, different methods of computing such "pure superspace integrals" are explored. As an example, a typical correlator encountered in the two-loop calculations of section 4 is considered:

$$
\begin{equation*}
F\left(k_{i}, u_{i}\right)=k_{a}^{2} k_{m}^{2} k_{p}^{3} k_{r}^{4}\left\langle\left(\lambda \gamma^{m n p q[r} \lambda\right)\left(\lambda \gamma^{s]} u_{1}\right)\left(\theta \gamma_{n}^{a b} \theta\right)\left(\theta \gamma_{b} u_{2}\right)\left(\theta \gamma_{q} u_{3}\right)\left(\theta \gamma_{s} u_{4}\right)\right\rangle \tag{2.3}
\end{equation*}
$$

Here, $k^{i}$ and $u_{i}$ are the momenta and spinor wavefunctions of the four external particles.

### 2.1 Symmetry considerations and tensorial formulae

One systematic approach to evaluate the zero mode integrals is to find expressions for all tensors that can be formed from (2.1). By Fierz transformations, one can always write the product of two $\theta$ spinors as $\left(\theta \gamma^{[3]} \theta\right)$, where $\gamma^{[k]}$ denotes the antisymmetrised product of $k$ gamma matrices. Due to the pure spinor constraint, the only bilinear in $\lambda$ is $\left(\lambda \gamma^{[5]} \lambda\right)$, and it is thus sufficient to consider the three cases

$$
\begin{equation*}
\left\langle\left(\lambda \gamma^{[5]} \lambda\right)\left(\lambda\left\{\gamma^{[1]} \text { or } \gamma^{[3]} \text { or } \gamma^{[5]}\right\} \theta\right)\left(\theta \gamma^{[3]} \theta\right)\left(\theta \gamma^{[3]} \theta\right)\right\rangle \tag{2.4}
\end{equation*}
$$

Lorentz invariance then implies that it must be possible to express these tensors as sums of suitably symmetrised products of metric tensors, resulting in a parity-even expression, plus a parity-odd part made up from terms which in addition contain an epsilon tensor. The parity-even parts may be constructed [6] starting from the most general ansatz compatible with the symmetries of the correlator and then using spinor identities along with the normalisation (2.2) to determine all coefficients in the ansatz. Duality properties of the spinor bilinears can be used to determine the parity-odd part [7]. An extensive (and almost exhaustive) list of correlators is found in [11], including the $\left(\lambda \gamma^{[1]} \theta\right)$ and $\left(\lambda \gamma^{[3]} \theta\right)$ cases of the above list:

$$
\begin{align*}
\left\langle\left(\lambda \gamma^{m n p q r} \lambda\right)\left(\lambda \gamma^{u} \theta\right)\left(\theta \gamma^{f g h} \theta\right)\left(\theta \gamma^{j k l} \theta\right)\right\rangle= & -\frac{4}{35}\left(\delta_{\bar{m} \bar{n} \bar{p} \bar{q} \bar{r}}^{m n p q r}+\frac{1}{5!} \varepsilon^{m n p q r}{ }_{\bar{m} \bar{n} \bar{p} \bar{q} \bar{r})}\right. \\
\times & {\left[\delta_{f g}^{\bar{m} \bar{n}} \delta_{j k}^{\bar{p} \bar{q}}\left(\delta_{l}^{\bar{r}} \delta_{u}^{h}+\delta_{h}^{\bar{r}} \delta_{u}^{l}-\delta_{u}^{\bar{r}} \delta_{l}^{h}\right)\right]_{[f g h][j k l]} }  \tag{2.5}\\
\left\langle\left(\lambda \gamma^{m n p q r} \lambda\right)\left(\lambda \gamma^{s t u} \theta\right)\left(\theta \gamma^{f g h} \theta\right)\left(\theta \gamma^{j k l} \theta\right)\right\rangle= & -\frac{24}{35}\left(\delta_{\bar{m} \bar{n} \bar{p} \bar{q} \bar{r}}^{m n p q r}+\frac{1}{5!} \varepsilon^{m n p q r}{ }_{\bar{m} \bar{n} \bar{p} \bar{q} \bar{r}}\right) \\
& \times\left[\delta_{j}^{\bar{m}} \delta_{f g}^{\bar{n} \bar{p}} \delta_{s}^{\bar{q}} \delta_{l}^{t}\left(\delta_{h}^{\bar{r}} \delta_{u}^{k}-\delta_{h}^{k} \delta_{u}^{\bar{r}}\right)\right]_{[f g h][j k l](f g h \leftrightarrow j k l)} \tag{2.6}
\end{align*}
$$

(Here, the brackets $(f g h \leftrightarrow j k l)$ denote symmetrisation under simultaneous interchange of $f g h$ with $i j k$, with weight one.) The remaining correlator with the $\left(\lambda \gamma^{[5]} \theta\right)$ factor can be derived in the same way, using an ansatz consisting of six parity-even structures. Taking a trace between the two $\gamma^{[5]}$ factors and noting that

$$
\eta_{a r}\left\langle\left(\lambda \gamma^{m n p q r} \lambda\right)\left(\lambda \gamma^{a b c d e} \theta\right) \ldots\right\rangle=-4\left\langle\left(\lambda \gamma^{m n p q[b} \lambda\right)\left(\lambda \gamma^{c d e]} \theta\right) \ldots\right\rangle
$$

one finds a relation to (2.6). This is sufficient to determine all coefficients in the ansatz, and the result is

$$
\begin{align*}
& \left\langle\left(\lambda \gamma^{m n p q r} \lambda\right)\left(\lambda \gamma^{a b c d e} \theta\right)\left(\theta \gamma^{f g h} \theta\right)\left(\theta \gamma^{j k l} \theta\right)\right\rangle=\frac{16}{7}\left(\delta_{\bar{m} \bar{n} \bar{p} \bar{q} \bar{r}}^{m n p q r}+\frac{1}{5!} \varepsilon^{m n p q r} \bar{m} \bar{n} \bar{p} \bar{q} \bar{r}\right) \\
& \quad \times\left[\delta_{a b c}^{\bar{m} \bar{n} \bar{p}} \delta_{j}^{f} \delta_{g}^{d} \delta_{k}^{\bar{q}}\left(-\delta_{h}^{e} \delta_{l}^{\bar{r}}+2 \delta_{l}^{e} \delta_{h}^{\bar{r}}\right)+\delta_{a b}^{\bar{m} \bar{n}} \delta_{f g}^{c d} \delta_{j k}^{\bar{p} \bar{q}}\left(\delta_{h}^{e} \delta_{l}^{\bar{r}}-3 \delta_{l}^{e} \delta_{h}^{\bar{r}}\right)\right]_{[a b c d e][f g h][j k l](f g h \leftrightarrow j k l)} \tag{2.7}
\end{align*}
$$

One may find it surprising that the derivation of these tensorial expressions only made use of properties of (pure) spinors, and of the normalisation condition (2.2). However, it can be seen from representation theory that the correlator (2.1) is uniquely characterised, up to normalisation, by its symmetry. To see this, note that 12] the spinor products $\lambda^{3}$ and $\theta^{5}$ transform in

$$
\begin{align*}
\lambda^{(\alpha} \lambda^{\beta} \lambda^{\gamma)}: & \operatorname{Sym}^{3} S^{+} & =[00003] \oplus[10001]  \tag{2.8}\\
\theta^{\left[\delta_{1}\right.} \ldots \theta^{\left.\delta_{5}\right]}: & \operatorname{Alt}^{5} S^{+} & =[00030] \oplus[11010] .
\end{align*}
$$

(Here, $\lambda$ and $\theta$ are taken to be in the $S^{+}$irrep of $\operatorname{SO}(1,9)$, with Dynkin label [00001].) The tensor product of these contains only one copy of the trivial representation. This applies to any spinors $\lambda$, which means that the pure spinor property cannot be essential to the derivation of the tensorial identities. The use of the pure spinor constraint merely allows for simpler derivations of the same identities.

As an illustration of this approach, consider the correlator of eq. (2.3). Leaving the momenta aside for the moment by setting $F=k_{a}^{2} k_{m}^{2} k_{p}^{3} k_{r}^{4} \tilde{F}$, the task is to compute

$$
\tilde{F}=\left\langle\left(\lambda \gamma^{m n p q[r} \lambda\right)\left(\lambda \gamma^{s]} u_{1}\right)\left(\theta \gamma_{n}{ }^{a b} \theta\right)\left(\theta \gamma_{b} u_{2}\right)\left(\theta \gamma_{q} u_{3}\right)\left(\theta \gamma_{s} u_{4}\right)\right\rangle .
$$

After applying two Fierz transformations,

$$
\begin{aligned}
\tilde{F}=( & \frac{1}{16}\left\langle\left(\lambda \gamma^{m n p q[r \mid} \lambda\right)\left(\lambda \gamma^{c} \theta\right)\left(\theta \gamma_{n}{ }^{a b} \theta\right)\left(\theta \gamma^{j k l} \theta\right)\right\rangle\left(u_{1} \gamma^{\mid s]} \gamma_{c} \gamma_{b} u_{2}\right) \\
& +\frac{1}{3!\cdot 16}\left\langle\left(\lambda \gamma^{m n p q[r \mid} \lambda\right)\left(\lambda \gamma^{c d e} \theta\right)\left(\theta \gamma_{n}{ }^{a b} \theta\right)\left(\theta \gamma^{j k l} \theta\right)\right\rangle\left(u_{1} \gamma^{\mid s]} \gamma_{c d e} \gamma_{b} u_{2}\right) \\
& \left.+\frac{1}{2 \cdot 5!\cdot 16}\left\langle\left(\lambda \gamma^{m n p q[r \mid} \lambda\right)\left(\lambda \gamma^{c d e f g} \theta\right)\left(\theta \gamma_{n}{ }^{a b} \theta\right)\left(\theta \gamma^{j k l} \theta\right)\right\rangle\left(u_{1} \gamma^{\mid s]} \gamma_{c d e f g} \gamma_{b} u_{2}\right)\right) \\
\times & \frac{1}{3!\cdot 16}\left(u_{3} \gamma_{q} \gamma_{j k l} \gamma_{s} u_{4}\right),
\end{aligned}
$$

one obtains a combination of the fundamental correlators listed in (2.1), (2.6) and (2.7). A reliable evaluation of the numerous index symmetrisations is made possible by the use of a computer algebra program. In doing these calculations with Mathematica, an essential tool is the GAMMA package [13], expanding the products of gamma matrices in a $\gamma^{[k]}$ basis. The result consists of two parts, $\tilde{F}=\tilde{F}^{(\delta)}+\tilde{F}^{(\varepsilon)}$, with

$$
\begin{align*}
\tilde{F}^{(\delta)}= & \frac{1}{560}\left(u_{1} \gamma^{m p r} u_{2}\right)\left(u_{3} \gamma^{a} u_{4}\right)+\frac{7}{720} \delta_{p}^{a} \delta_{r}^{m}\left(u_{1} \gamma^{i} u_{2}\right)\left(u_{3} \gamma_{i} u_{4}\right)+\ldots \\
& -\frac{1}{1680}\left(u_{1} \gamma^{a i_{1} i_{2}} u_{2}\right)\left(u_{3} \gamma^{m p r}{ }_{i_{1} i_{2} u_{4}}\right) \quad(92 \text { terms })  \tag{2.9}\\
\tilde{F}^{(\varepsilon)}= & -\frac{1}{1209600} \varepsilon_{i_{1} \ldots i_{7}}{ }^{m p r}\left(u_{1} \gamma^{i_{1} \ldots i_{7}} u_{2}\right)\left(u_{3} \gamma^{a} u_{4}\right)+\ldots \\
& -\frac{1}{604800} \varepsilon^{a m p r}{ }_{i_{1} \ldots i_{6}}\left(u_{1} \gamma^{i_{3} \ldots i_{9}} u_{2}\right)\left(u_{3} \gamma^{i_{1} i_{2}}{ }_{i_{7} i_{8} i_{9}} u_{4}\right) \quad(34 \text { terms }) \tag{2.10}
\end{align*}
$$

The epsilon tensors in the second part can be eliminated using the fact that the $u_{i}$ are chiral spinors: If all the indices on $\gamma^{[k]} u_{i}$ are contracted into an epsilon tensor, one uses

$$
\begin{equation*}
\varepsilon_{i_{1} \ldots i_{k^{\prime}} j_{1} \ldots j_{k}} \gamma^{j_{1} \ldots j_{k}} \gamma^{11}=(-)^{\frac{1}{2} k(k+1)} k!\gamma_{i_{1} \ldots i_{k^{\prime}}}, \tag{2.11}
\end{equation*}
$$

where $\gamma^{11}=\frac{1}{10!} \varepsilon_{i_{0} \ldots i_{9}} \gamma^{i_{0} \ldots i_{9}}$. More generally, if all but $r$ indices of $\gamma^{[k]} u_{i}$ are contracted,

$$
\begin{equation*}
\varepsilon_{i_{1} \ldots i_{k^{\prime}} j_{1} \ldots j_{k}} \gamma^{p_{1} \ldots p_{r} j_{1} \ldots j_{k}} \gamma^{11}=(-)^{\frac{1}{2} k(k+1)} k!\frac{k^{\prime}!}{\left(k^{\prime}-r\right)!} \delta_{\left[i_{1} \ldots i_{r}\right.}^{p_{r} \ldots p_{1}} \gamma_{\left.i_{r+1} \ldots i_{k}^{\prime}\right]} . \tag{2.12}
\end{equation*}
$$

The result of these manipulations is

$$
\begin{align*}
\tilde{F}^{(\varepsilon)}= & -\frac{1}{560}\left(u_{1} \gamma^{m p r} u_{2}\right)\left(u_{3} \gamma^{a} u_{4}\right)-\frac{1}{280} \delta_{r}^{p}\left(u_{1} \gamma^{a m i} u_{2}\right)\left(u_{3} \gamma_{i} u_{4}\right)+\ldots \\
& +\frac{9}{11200}\left(u_{1} \gamma^{i_{1} i_{2} i_{3}} u_{2}\right)\left(u_{3} \gamma^{a m p r}{ }_{i_{1} i_{2} i_{3}} u_{4}\right) \quad(53 \text { terms }) \tag{2.13}
\end{align*}
$$

(Note that while the epsilon terms in the basic correlator formulae were easily obtained from the delta terms by using Poincaré duality, this cannot be done here in any obvious way.) The last step in the evaluation of (2.3) is to contract with the momenta, $F=k_{a}^{2} k_{m}^{2} k_{p}^{3} k_{r}^{4} \tilde{F}$, and to simplify the expressions using the on-shell identities $\sum_{i} k_{i}=0, k_{i}^{2}=0, \not k_{i} u_{i}=0$. It is shown in appendix $\widehat{A .2}$ that there are only ten independent scalars, denoted by $B_{1} \ldots B_{10}$, that can be formed from four momenta and the four spinors $u_{1} \ldots u_{4}$. With respect to this basis, the result is

$$
\begin{align*}
F^{(\delta)} & =\frac{1}{48 \cdot 10080}\left(695 s_{12}\left(u_{1} \not k_{3} u_{2}\right)\left(u_{3} k_{1} u_{4}\right)+\cdots+233 s_{13}^{2}\left(u_{1} \gamma^{a} u_{2}\right)\left(u_{3} \gamma_{a} u_{4}\right)\right)  \tag{7terms}\\
& =\frac{1}{48 \cdot 10080}(695,775,0,-80,356,356,0,233,233,0)_{B_{1} \ldots B_{10}}, \\
F^{(\varepsilon)} & =\frac{1}{48 \cdot 10080}(-23,-7,0,-16,28,28,0,7,7,0)_{B_{1} \ldots B_{10}}, \\
F & =\frac{1}{10080}(14,16,0,-2,8,8,0,5,5,0)_{B_{1} \ldots B_{10}}, \tag{2.14}
\end{align*}
$$

where $s_{i j}=k_{i} \cdot k_{j}$.

### 2.2 A spinorial formula

While the derivation of tensorial identities for correlators of the form (2.4) is relatively straightforward and elegant, it may be a tedious task to transform the expressions encountered in amplitude calculations to match this pattern. As seen in the example calculated above, this is particularly true if additional spinors are involved, making it necessary to apply Fierz transformations. It is therefore desirable to use a covariant correlator expression with open spinor indices. Such an expression was given in [1], 2]:

$$
\begin{equation*}
\bar{T}^{\alpha \beta \gamma, \delta_{1} \ldots \delta_{5}}=N^{-1}\left[\left(\gamma^{m}\right)^{\alpha \delta_{1}}\left(\gamma^{n}\right)^{\beta \delta_{2}}\left(\gamma^{p}\right)^{\gamma \delta_{3}}\left(\gamma_{m n p}\right)^{\delta_{4} \delta_{5}}\right]_{(\alpha \beta \gamma)\left[\delta_{1} \ldots \delta_{5}\right]}, \tag{2.15}
\end{equation*}
$$

where $N$ is a normalisation constant and the brackets ()] denote (anti-)symmetrisation with weight one. (Note that the right hand side is automatically gamma-matrix traceless: any gamma-trace

$$
\left(\gamma^{r}\right)_{\alpha \beta} \times\left(\gamma^{m}\right)^{\alpha\left[\delta_{1} \mid\right.}\left(\gamma^{n}\right)^{\beta\left|\delta_{2}\right|}\left(\gamma^{p}\right)^{\gamma \mid \delta_{3}}\left(\gamma_{m n p}\right)^{\left.\delta_{4} \delta_{5}\right]}=-\left(\gamma^{m n r}\right)^{\left[\delta_{1} \delta_{2}\right.}\left(\gamma_{m n p}\right)^{\delta_{3} \delta_{4}}\left(\gamma^{p}\right)^{\left.\delta_{5}\right] \gamma}=0
$$

vanishes due to the double-trace identity $\left(\gamma_{a b} \theta\right)^{\alpha}\left(\theta \gamma^{a b c} \theta\right)=0$, which follows from the fact that the tensor product $\left(\mathrm{Alt}^{3} S^{+}\right) \otimes S^{-}$does not contain a vector representation and
therefore the vector $\left(\psi \gamma_{a b} \theta\right)\left(\theta \gamma^{a b c} \theta\right)$ has to vanish for all spinors $\psi$, and can also be shown by applying a Fierz transformation.) This prescription was originally motivated by the fermionic expansion of the Yang-Mills antighost vertex operator $V$,

$$
\begin{equation*}
V=T_{\alpha \beta \gamma, \delta_{1} \ldots \delta_{5}} \lambda^{\alpha} \lambda^{\beta} \lambda^{\gamma} \theta^{\delta_{1}} \ldots \theta^{\delta_{5}} \tag{2.16}
\end{equation*}
$$

with

$$
T_{\alpha \beta \gamma, \delta_{1} \ldots \delta_{5}}=\left[\left(\gamma^{m}\right)_{\alpha \delta_{1}}\left(\gamma^{n}\right)_{\beta \delta_{2}}\left(\gamma^{p}\right)_{\gamma \delta_{3}}\left(\gamma_{m n p}\right)_{\delta_{4} \delta_{5}}\right]_{(\alpha \beta \gamma)\left[\delta_{1} \ldots \delta_{5}\right]},
$$

where $T$ is related to $\bar{T}$ by a parity transformation, up to the overall constant $N$. (Seeing that $\bar{T}$ is uniquely determined by its symmetries, any covariant expression will be proportional to $\bar{T}$, after symmetrisation of the spinor indices, and this is merely the simplest choice.)

Equation (2.15) immediately yields an algorithm to convert any correlator into traces of gamma matrices or, if additional spinors are involved, bilinears in those spinors. It is, however, already very tiresome to determine the normalisation constant $N$ by hand. The main advantage of this approach is that it clearly lends itself to implementation on a computer algebra system, which can easily carry out the spinor index symmetrisations, simplify the gamma products (again using the GAMMA package), and compute the traces. For example,

$$
\begin{aligned}
N\langle V\rangle= & {\left[\left(\gamma^{m}\right)^{\alpha \delta_{1}}\left(\gamma^{n}\right)^{\beta \delta_{2}}\left(\gamma^{p}\right)^{\gamma \delta_{3}}\left(\gamma_{m n p}\right)^{\delta_{4} \delta_{5}}\right]_{(\alpha \beta \gamma)\left[\delta_{1} \ldots \delta_{5}\right]}\left(\gamma_{x}\right)_{\alpha \delta_{1}}\left(\gamma_{y}\right)_{\beta \delta_{2}}\left(\gamma_{z}\right)_{\gamma \delta_{3}}\left(\gamma^{x y z}\right)_{\delta_{4} \delta_{5}} } \\
= & -\frac{1}{60} \operatorname{Tr}\left(\gamma_{x} \gamma^{m}\right) \operatorname{Tr}\left(\gamma_{y} \gamma^{n}\right) \operatorname{Tr}\left(\gamma_{z} \gamma^{p}\right) \operatorname{Tr}\left(\gamma^{x y z} \gamma_{p n m}\right)+\ldots \\
& -\frac{1}{60} \operatorname{Tr}\left(\gamma_{z} \gamma_{p n m} \gamma^{z y x} \gamma^{n} \gamma_{x} \gamma^{m} \gamma_{y} \gamma^{p}\right) \quad(60 \text { terms }) \\
= & 5160960 .
\end{aligned}
$$

The correct normalisation is therefore obtained by setting $N=5160960$.
Returning to the example correlator (2.3), one finds that the calculation is by far simpler than with the previous method. After carrying out the symmetrisations $(\alpha \beta \gamma)\left[\delta_{i}\right]$, one obtains

$$
\begin{align*}
N \tilde{F}= & \frac{1}{60} \operatorname{Tr}\left(\gamma_{x} \gamma^{a b}{ }_{n} \gamma_{y} \gamma^{m n p q[r \mid}\right)\left(u_{3} \gamma_{q} \gamma^{x y z} \gamma_{s} u_{4}\right)\left(u_{1} \gamma^{\mid s]} \gamma_{z} \gamma_{b} u_{2}\right)+\ldots \\
& -\frac{1}{30}\left(u_{2} \gamma_{b} \gamma^{x y z} \gamma_{q} u_{3}\right)\left(u_{1} \gamma_{s} \gamma_{y} \gamma^{a b}{ }_{n} \gamma_{x} \gamma^{m n p q[r} \gamma_{z} \gamma^{s]} u_{4}\right), \tag{24terms}
\end{align*}
$$

where elementary index re-sorting has reduced the number of terms from 60 to 24 . Expanding the gamma products leads to
$N \tilde{F}=\frac{476}{5} \delta_{r}^{p}\left(u_{1} \gamma^{m} u_{4}\right)\left(u_{2} \gamma^{a} u_{3}\right)+\cdots+\frac{8}{15}\left(u_{1} \gamma^{a i_{1} i_{2} i_{3} i_{4}} u_{2}\right)\left(u_{3} \gamma^{m p r}{ }_{i_{1} i_{2} i_{3} i_{4}} u_{4}\right)$,
(294 terms)
which, in contrast to (2.10), contains no epsilon terms as there are not enough free indices present. Note that this intermediate result contains terms with with $u_{1}$ paired with $u_{3}$ or $u_{4}$, so it is not possible to directly compare to eqs. (2.9) and (2.13). However, after contracting
with the momenta $k_{a}^{2} k_{m}^{2} k_{p}^{3} k_{r}^{4}$ and decomposing the result in the basis $B_{1} \ldots B_{10}$, one again obtains

$$
\begin{equation*}
F=\frac{1}{10080}(14,16,0,-2,8,8,0,5,5,0)_{B_{1} \ldots B_{10}}, \tag{2.17}
\end{equation*}
$$

in agreement with (2.14).
The algorithm just outlined will be the method of choice for all correlator calculations in the later sections of this paper and can easily be applied to a wider range of problems. The only limitation is that the larger the number of gamma matrices and open indices of the correlator, the slower the computer evaluation will be. For example, the correlator considered in eq. (5.2) of [11],

$$
\begin{align*}
t_{10}^{m n m_{1} n_{1} \ldots m_{4} n_{4}} & \equiv\left\langle\left(\lambda \gamma^{p} \gamma^{m_{1} n_{1}} \theta\right)\left(\lambda \gamma^{q} \gamma^{m_{2} n_{2}} \theta\right)\left(\lambda \gamma^{r} \gamma^{m_{3} n_{3}} \theta\right)\left(\theta \gamma^{m} \gamma^{n} \gamma_{p q r} \gamma^{m_{4} n_{4}} \theta\right)\right\rangle  \tag{2.18}\\
& =-\frac{2}{45}\left(\eta^{m n} t_{8}^{m_{1} n_{1} \ldots m_{4} n_{4}}-\frac{1}{2} \varepsilon^{m n m_{1} n_{1} \ldots m_{4} n_{4}}\right) \tag{2.19}
\end{align*}
$$

can still be verified with this method but this already requires substantial runtime.

### 2.3 Component-based approach

A third method to evaluate the zero mode integrals consists of choosing a gamma matrix representation, expanding the integrand as a polynomial in spinor components, and then applying (2.15) to the individual monomials. This procedure seems particularly appealing if at some stage of the calculation one works with a matrix representation anyhow, in order to reduce the results to a canonical form (e.g. as outlined in appendix A). An efficient decomposition algorithm (of $k^{4} u_{1} u_{2} u_{3} u_{4}$ scalars, say) only needs a few non-zero momentum and spinor wavefunction components to distinguish all independent scalars, and therefore $k$ and $u$ can be replaced by sparse vectors. Furthermore, a trivial observation allows for a much quicker numeric evaluation of correlator components than a naive use of (2.15): In view of (2.16), one can equivalently compute the components of the paritytransformed expression $\bar{V}=\left(\bar{\lambda} \gamma^{m} \bar{\theta}\right)\left(\bar{\lambda} \gamma^{n} \bar{\theta}\right)\left(\bar{\lambda} \gamma^{p} \bar{\theta}\right)\left(\bar{\theta} \gamma_{m n p} \bar{\theta}\right)$, where $\bar{\lambda}$ and $\bar{\theta}$ are spinors of chirality opposite to that of $\lambda, \theta$. In the representation given in appendix B, $\bar{V}$ coincides with $\left.V\right|_{\lambda \rightarrow \bar{\lambda}, \theta \rightarrow \bar{\theta}}$, and

$$
\begin{equation*}
V=192 \lambda^{9} \lambda^{9} \lambda^{9} \theta^{1} \theta^{2} \theta^{3} \theta^{4} \theta^{9}+\cdots+480 \lambda^{1} \lambda^{2} \lambda^{3} \theta^{1} \theta^{9} \theta^{10} \theta^{13} \theta^{15}+\ldots \tag{100352terms}
\end{equation*}
$$

The monomials in the fermionic expansion of $\bar{V}$ then correspond to the arguments of non-zero correlators, and the coefficients of those monomials are, up to normalisation and symmetry factors, the correlator values.

Unfortunately, it turns out that the complexity of typical correlators (e.g. the one given in (2.3)) makes it difficult to carry out the expansion in fermionic components in any straightforward way and limits this method to special applications. For example, the coefficients in (2.18) can be checked relatively easily by choosing particular index values, such as

$$
\begin{align*}
\left\langle\left(\lambda \gamma^{p} \gamma^{12} \theta\right)\left(\lambda \gamma^{q} \gamma^{21} \theta\right)\left(\lambda \gamma^{r} \gamma^{34} \theta\right)\left(\theta \gamma^{0} \gamma^{0} \gamma_{p q r} \gamma^{43} \theta\right)\right\rangle= & \left\langle 12 \lambda^{1} \lambda^{1} \lambda^{1} \theta^{1} \theta^{9} \theta^{10} \theta^{11} \theta^{12}+\ldots\right. \\
& \left.+12 \lambda^{16} \lambda^{16} \lambda^{16} \theta^{5} \theta^{6} \theta^{7} \theta^{8} \theta^{16}\right\rangle=\frac{1}{45} . \tag{2.20}
\end{align*}
$$

(For fixed values of $p q r$, one gets no more than about $10^{5}$ monomials of the form $\lambda^{3} \theta^{5}$ ). This approach may thus still be helpful in situations where the result has been narrowed down to a simple ansatz.

## 3. One-loop amplitudes

The amplitude for the scattering of four massless states of the type IIB superstring was computed [2] in the pure spinor formalism as

$$
\begin{equation*}
\mathcal{A}=K \bar{K} \int \frac{d^{2} \tau}{(\operatorname{Im} \tau)^{5}} \int d^{2} z_{2} \int d^{2} z_{3} \int d^{2} z_{4} \prod_{i<j} G\left(z_{i}, z_{j}\right)^{k_{i} \cdot k_{j}}, \tag{3.1}
\end{equation*}
$$

where $G\left(z_{i}, z_{j}\right)$ is the scalar Green's function, and the kinematic factor is given by the product $K \bar{K}$ of left- and right-moving open superstring expressions,

$$
\begin{equation*}
K_{1 \text {-loop }}=\left\langle\left(\lambda A_{1}\right)\left(\lambda \gamma^{m} W_{2}\right)\left(\lambda \gamma^{n} W_{3}\right) \mathcal{F}_{4, m n}\right\rangle+(\operatorname{cycl}(234)) . \tag{3.2}
\end{equation*}
$$

Here the indices $1 \ldots 4$ label the external states and $" \cdots+(\operatorname{cycl}(234))$ " denotes the addition of two other terms obtained by cyclic permutation of the indices 234 . The spinor superfield $A_{\alpha}$ and its supercovariant derivatives, the vector gauge superfield $A_{m}=\frac{1}{8} \gamma_{m}^{\alpha \beta} D_{\alpha} A_{\beta}$ as well as the spinor and vector field strengths $W^{\alpha}=\frac{1}{10}\left(\gamma^{m}\right)^{\alpha \beta}\left(D_{\beta} A_{m}-\partial_{m} A_{\beta}\right)$ and $\mathcal{F}_{m n}=\frac{1}{8}\left(\gamma_{m n}\right)^{\alpha}{ }_{\beta} D_{\alpha} W^{\beta}=2 \partial_{[m} A_{n]}$, describe ten-dimensional super-Yang-Mills theory. The physical fields of this theory, a gauge boson and a gaugino, are found in the leading components $A_{m} \mid=\zeta_{m}$ and $W^{\alpha} \mid=u^{\alpha}$ and correspond to the Neveu-Schwarz and Ramond superstring states.

Choosing a gauge where $\theta^{\alpha} A_{\alpha}=0$, the on-shell identities

$$
2 D_{(\alpha} A_{\beta)}=\gamma_{\alpha \beta}^{m} A_{m}, \quad D_{\alpha} W^{\beta}=\frac{1}{4}\left(\gamma^{m n}\right)_{\alpha}^{\beta} \mathcal{F}_{m n}
$$

have been used to derive recursive relations [10, 14, 15] for the fermionic expansion
$A_{\alpha}^{(n)}=\frac{1}{n+1}\left(\gamma^{m} \theta\right)_{\alpha} A_{m}^{(n-1)}, \quad A_{m}^{(n)}=\frac{1}{n}\left(\theta \gamma_{m} W^{(n-1)}\right), \quad W^{\alpha(n)}=-\frac{1}{2 n}\left(\gamma^{m n} \theta\right)^{\alpha} \partial_{m} A_{n}^{(n-1)}$,
where $\left.f^{(n)}=\frac{1}{n!} \theta^{\alpha_{n}} \cdots \theta^{\alpha_{1}}\left(D_{\alpha_{1}} \cdots D_{\alpha_{n}} f\right) \right\rvert\,$. These recursion relations were explicitly solved in 10], reducing the fermionic expansion to a simple repeated application of the derivative operator $\mathcal{O}_{m}{ }^{q}=\frac{1}{2}\left(\theta \gamma_{m}{ }^{q p} \theta\right) \partial_{p}$ :

$$
\begin{align*}
A_{m}^{(2 k)} & =\frac{1}{(2 k)!}\left[\mathcal{O}^{k}\right]_{m}{ }^{q} \zeta_{q}, \\
A_{m}^{(2 k+1)} & =\frac{1}{(2 k+1)!}\left[\mathcal{O}^{k}\right]_{m}{ }^{q}\left(\theta \gamma_{q} u\right) . \tag{3.3}
\end{align*}
$$

With this solution at hand, one has all ingredients to evaluate the kinematic factor (3.2) for the three cases of zero, two, or four fermionic states.

### 3.1 Review: four bosons

The kinematic factor involving four bosons was considered in (7) and this calculation will now be reviewed briefly. First, note that the outcome is not fixed by symmetry: The result must be gauge invariant [2] and therefore expressible in terms of the field strengths $F_{1} \ldots F_{4}$. The cyclic symmetrisation in (3.2) yields expressions symmetric in $F_{2}, F_{3}, F_{4}$, and acting on scalars constructed from the $F_{i}$ only, the (234) symmetrisation is equivalent to complete symmetrisation in all labels (1234). Thus the result must be a linear combination of the two gauge invariant symmetric $F^{4}$ scalars, namely the single trace $\operatorname{Tr}\left(F_{(1} F_{2} F_{3} F_{4)}\right)$ and double trace $\operatorname{Tr}\left(F_{(1} F_{2}\right) \operatorname{Tr}\left(F_{3} F_{4)}\right)$, leaving one relative coefficient to be determined.

Since all four states are of the same kind, one may first evaluate the correlator for one labelling and then carry out the cyclic symmetrisation:

$$
K_{1-\mathrm{loop}}^{(4 \mathrm{~B})}=\left.\left\langle\left(\lambda A_{1}\right)\left(\lambda \gamma^{m} W_{2}\right)\left(\lambda \gamma^{n} W_{3}\right) \mathcal{F}_{4, m n}\right\rangle\right|_{4 \mathrm{~B}}+(\operatorname{cycl}(234))
$$

The different ways to saturate $\theta^{5}$ result in a sum of terms of the form

$$
\begin{equation*}
X_{A B C D}=\left\langle\left(\lambda A_{1}^{(A)}\right)\left(\lambda \gamma^{m} W_{2}^{(B)}\right)\left(\lambda \gamma^{n} W_{3}^{(C)}\right) \mathcal{F}_{4, m n}^{(D)}\right\rangle \tag{3.4}
\end{equation*}
$$

with $A+B+C+D=5$ and $A, B, C$ odd, $D$ even:

$$
\left.\left\langle\left(\lambda A_{1}\right)\left(\lambda \gamma^{m} W_{2}\right)\left(\lambda \gamma^{n} W_{3}\right) \mathcal{F}_{4, m n}\right\rangle\right|_{4 \mathrm{~B}}=X_{3110}+X_{1310}+X_{1130}+X_{1112}
$$

Note that $X_{1130}+(\operatorname{cycl}(234))=X_{1310}+(\operatorname{cycl}(234))$, so only one of those two terms needs to be computed. Using (3.3) for the superfield expansions, one obtains

$$
\begin{array}{ll}
X_{3110}=\frac{15}{64} F_{m n}^{1} F_{p q}^{2} F_{r s}^{3} F_{t u}^{4} \tilde{X}_{3110}, & \tilde{X}_{3110}=\left\langle\left(\lambda \gamma^{[t \mid} \gamma^{p q} \theta\right)\left(\lambda \gamma^{\mid u]} \gamma^{r s} \theta\right)\left(\lambda \gamma_{a} \theta\right)\left(\theta \gamma^{a m n} \theta\right)\right\rangle, \\
X_{1112}=\frac{15}{16} k_{m}^{4} \zeta_{n}^{1} F_{p q}^{2} F_{r s}^{3} F_{t u}^{4} \tilde{X}_{1112}, & \tilde{X}_{1112}=\left\langle\left(\lambda \gamma^{[m \mid} \gamma^{p q} \theta\right)\left(\lambda \gamma^{\mid a]} \gamma^{r s} \theta\right)\left(\lambda \gamma^{n} \theta\right)\left(\theta \gamma_{a}^{t u} \theta\right)\right\rangle, \\
X_{1310}=\frac{5}{16} k_{m}^{3} \zeta_{n}^{1} F_{p q}^{2} F_{r s}^{3} F_{t u}^{4} \tilde{X}_{1310}, & \tilde{X}_{1310}=\left\langle\left(\lambda \gamma^{[t \mid} \gamma^{m a} \theta\right)\left(\lambda \gamma^{\mid u]} \gamma^{r s} \theta\right)\left(\lambda \gamma^{n} \theta\right)\left(\theta \gamma_{a}^{p q} \theta\right)\right\rangle .
\end{array}
$$

The method outlined in section 2.2 is readily applicable to these correlators. For example, for $X_{3111}$, the trace evaluation yields

$$
\begin{aligned}
& \tilde{X}_{3110}= N^{-1}\left[\frac{1}{60} \operatorname{Tr}\left(\gamma_{a} \gamma^{z}\right) \operatorname{Tr}\left(\gamma_{x y z} \gamma^{a n m}\right) \operatorname{Tr}\left(\gamma^{x} \gamma_{q p} \gamma^{[t \mid}\right) \operatorname{Tr}\left(\gamma^{y} \gamma_{s r} \gamma^{\mid u]}\right)+\cdots\right. \\
&\left.\cdots+\frac{1}{60} \operatorname{Tr}\left(\gamma^{[u \mid} \gamma_{r s} \gamma_{z y x} \gamma_{q p} \gamma^{\mid t]} \gamma^{x} \gamma_{a} \gamma^{y} \gamma^{m n a} \gamma^{z}\right)\right] \quad(60 \text { terms }) \\
&=\left(\frac{2}{35} \delta_{r s}^{m p} \delta_{t u}^{n q}-\frac{1}{315} \delta_{t u}^{m n} \delta_{r s}^{p q}-\frac{1}{45} \delta_{r s}^{m n} \delta_{t u}^{p q}+\frac{26}{315} \delta_{p r}^{m n} \delta_{t u}^{q s}\right)_{[m n][p q][r s][t u](p q \leftrightarrow r s)}
\end{aligned}
$$

Upon contracting with the field strengths, momenta and polarisations, and symmetrising over the cyclic permutations (234) (with weight 3), one finds that all three contributions are separately gauge invariant:

$$
\begin{aligned}
X_{3110}+(\operatorname{cycl}(234)) & =\frac{11}{112} \operatorname{Tr}\left(F_{(1} F_{2} F_{3} F_{4)}\right)-\frac{1}{56} \operatorname{Tr}\left(F_{(1} F_{2}\right) \operatorname{Tr}\left(F_{3} F_{4)}\right) \\
X_{1112}+(\operatorname{cycl}(234)) & =\frac{19}{448} \operatorname{Tr}\left(F_{(1} F_{2} F_{3} F_{4)}\right)-\frac{31}{1792} \operatorname{Tr}\left(F_{(1} F_{2}\right) \operatorname{Tr}\left(F_{3} F_{4)}\right) \\
X_{1310}+(\operatorname{cycl}(234)) & =\frac{3}{512}\left(4 \operatorname{Tr}\left(F_{(1} F_{2} F_{3} F_{4)}\right)-\operatorname{Tr}\left(F_{(1} F_{2}\right) \operatorname{Tr}\left(F_{3} F_{4)}\right)\right)
\end{aligned}
$$

The sum $X_{3110}+X_{1112}$ has the right ratio of single- and double-trace terms to be proportional to the well-known result $t_{8} F^{4}$, and the last line exhibits the right ratio by itself. The overall kinematic factor is therefore

$$
\begin{equation*}
K_{1-\mathrm{loop}}^{4 \mathrm{~B}}=\frac{3}{64}\left(4 \operatorname{Tr}\left(F_{(1} F_{2} F_{3} F_{4)}\right)-\operatorname{Tr}\left(F_{(1} F_{2}\right) \operatorname{Tr}\left(F_{3} F_{4}\right)\right)=\frac{1}{128} t_{8} F^{4}, \tag{3.5}
\end{equation*}
$$

in agreement with the expressions derived in the RNS [16] and Green-Schwarz [17] formalisms.

### 3.2 Four fermions

The four-fermion kinematic factor could be evaluated in the same way as in the four-boson case by summing up all terms $X_{A B C D}, A+B+C+D=5$, now with $A, B, C$ even and $D$ odd. Note however that this time, the outcome is fixed by symmetry: The cyclic symmetrisation in (3.2) leads to a completely antisymmetric dependence on $u_{2}, u_{3}, u_{4}$. Acting on scalars of the form $k^{2} u_{1} u_{2} u_{3} u_{4}$, antisymmetrising over [234] is equivalent to antisymmetrising over [1234], and there is only one completely antisymmetric $k^{2} u_{1} u_{2} u_{3} u_{4}$ scalar. Without further calculation, one can infer that the kinematic factor is proportional to that scalar,

$$
K_{1-\text { loop }}^{4 \mathrm{~F}}=\text { const } \cdot\left(\left(u_{1} \not k_{3} u_{2}\right)\left(u_{3} \not k_{1} u_{4}\right)-\left(u_{1} \not k_{2} u_{3}\right)\left(u_{2} \not k_{1} u_{4}\right)+\left(u_{1} \not k_{2} u_{4}\right)\left(u_{2} \not k_{1} u_{3}\right)\right),
$$

which of course agrees with the RNS amplitude (see e.g. [16], eq. (3.67)).

### 3.3 Two bosons, two fermions

In evaluating (3.2) for two bosons and two fermions, the cyclic symmetrisations affect whether the $W$ and $\mathcal{F}$ superfields contribute bosons or fermions. Only the label of the $A_{\alpha}$ superfield stays unaffected, and one has to choose whether it should contribute a boson or a fermion. Since its fermionic expansion starts with the bosonic polarisation vector, $A_{1, \alpha} \sim\left(\$_{1} \theta\right)_{\alpha}$, the calculation can be simplified by choosing a labelling where particle 1 is a fermion. (Of course, the final result must be independent of this choice.) The assignment of the other three labels is then irrelevant and will be chosen as $f_{1} f_{2} b_{3} b_{4}$. Writing out the cyclic permutations, two of the three terms are essentially the same because they are related by interchange of the labels 3 and 4 . This interchange can be carried out after computing the correlator, an operation which will be denoted by $\pi_{34}$. The kinematic factor is then

$$
\begin{aligned}
K_{1-\text { loop }}^{2 \mathrm{B2F}}\left(f_{1} f_{2} b_{3} b_{4}\right)= & \left(1-\pi_{34}\right)\left\langle\left(\lambda A_{1}^{\text {(even })}\right)\left(\lambda \gamma^{m} W_{2}^{(\text {even })}\right)\left(\lambda \gamma^{n} W_{3}^{\text {(odd) })}\right) \mathcal{F}_{4, m n}^{(\text {even })}\right\rangle \\
& +\left\langle\left(\lambda A_{1}^{\text {(even })}\right)\left(\lambda \gamma^{m} W_{3}^{\text {(odd) })}\right)\left(\lambda \gamma^{n} W_{4}^{\text {(odd) })}\right) \mathcal{F}_{2, m n}^{\text {(odd) }}\right\rangle .
\end{aligned}
$$

While one expects the final result to be symmetric under exchange of 3 and 4, this is only the case for the second line of this expression (note that the exchange of the two $\mathrm{W}^{\text {(odd) }}$ factors gives a sign from interchange of odd numbers of $\theta$ ). The first line, on the other hand, is manifestly antisymmetric and should therefore give zero. Even then, unlike in the four-fermion calculation, the result is not fixed by symmetry. There are five independent
$k F_{1} F_{2} u_{3} u_{4}$ scalars (see appendix A, eq. (A.6)), denoted by $C_{1} \ldots C_{5}$, and there are two independent combinations of these scalars with the required [12](34) symmetry.

Expanding the superfields and collecting terms with $\theta^{5}$, the first line yields a combination of terms $X_{A B C D}$ with $A, B, D$ odd and $C$ even. There is only one $\theta^{5}$ combination coming from the second line, which will be denoted by $X_{2111}^{\prime} \equiv\left(-\pi_{24}\right) X_{2111}$ :

$$
K_{1-\text { loop }}^{2 \text { 2F2F }}=\left(1-\pi_{34}\right)\left(X_{4010}+X_{2210}+X_{2030}+X_{2012}\right)+X_{2111}^{\prime}
$$

with the correlators

$$
\begin{array}{ll}
X_{4010}=-\frac{1}{60} k_{q}^{1} k_{b}^{3} \zeta_{c}^{3} k_{m}^{4} \zeta_{n}^{4} \tilde{X}_{4010}, & \tilde{X}_{4010}=\left\langle\left(\lambda \gamma^{a} \theta\right)\left(\theta \gamma_{a}{ }^{p q} \theta\right)\left(\theta \gamma_{p} u_{1}\right)\left(\lambda \gamma^{[m} u_{2}\right)\left(\lambda \gamma^{n]} \gamma^{b c} \theta\right)\right\rangle \\
X_{2210}=\frac{1}{12} k_{b}^{2} k_{d}^{3} \zeta_{e}^{3} k_{m}^{4} \zeta_{n}^{4} \tilde{X}_{2210}, & \tilde{X}_{2210}=\left\langle\left(\lambda \gamma^{a} \theta\right)\left(\theta \gamma_{a} u_{1}\right)\left(\lambda \gamma^{[m \mid} \gamma^{b c} \theta\right)\left(\theta \gamma_{c} u_{2}\right)\left(\lambda \gamma^{\mid n]} \gamma^{d e} \theta\right)\right\rangle \\
X_{2030}=\frac{1}{36} k_{b}^{3} k_{d}^{3} \zeta_{e}^{3} k_{m}^{4} \zeta_{n}^{4} \tilde{X}_{2030}, & \tilde{X}_{2030}=\left\langle\left(\lambda \gamma^{a} \theta\right)\left(\theta \gamma_{a} u_{1}\right)\left(\lambda \gamma^{[m} u_{2}\right)\left(\lambda \gamma^{n]} \gamma^{b c} \theta\right)\left(\theta \gamma_{c}^{d e} \theta\right)\right\rangle \\
X_{2012}=\frac{1}{12} k_{b}^{3} \zeta_{c}^{3} k_{m}^{4} k_{d}^{4} \zeta_{e}^{4} \tilde{X}_{2012}, & \tilde{X}_{2012}=\left\langle\left(\lambda \gamma^{a} \theta\right)\left(\theta \gamma_{a} u_{1}\right)\left(\lambda \gamma^{[m} u_{2}\right)\left(\lambda \gamma^{n]} \gamma^{b c} \theta\right)\left(\theta \gamma_{n}^{d e} \theta\right)\right\rangle \\
X_{2111}^{\prime}=\frac{1}{6} k_{b}^{3} \zeta_{c}^{3} k_{d}^{4} \zeta_{e}^{4} k_{m}^{2} \tilde{X}_{2111}^{\prime}, & \tilde{X}_{2111}^{\prime}=\left\langle\left(\lambda \gamma^{a} \theta\right)\left(\theta \gamma_{a} u_{1}\right)\left(\lambda \gamma^{[m \mid} \gamma^{b c} \theta\right)\left(\lambda \gamma^{\mid n]} \gamma^{d e} \theta\right)\left(\theta \gamma_{n} u_{2}\right)\right\rangle
\end{array}
$$

Evaluating these expressions as outlined in section 2.2, the spinor wavefunctions $u_{i}$ present no complication. The part with the correct symmetry also takes the simplest form: One finds

$$
\left\langle\left(\lambda \gamma^{a} \theta\right)\left(\theta \gamma_{a} u_{1}\right)\left(\lambda \gamma^{m} \gamma^{b c} \theta\right)\left(\lambda \gamma^{n} \gamma^{d e} \theta\right)\left(\theta \gamma_{n} u_{2}\right)\right\rangle=-\frac{1}{240}\left(2 \delta_{m[d}^{b c}\left(u_{1} \gamma_{e]} u_{2}\right)+\delta_{m}^{[b}\left(u_{1} \gamma^{c] d e} u_{2}\right)\right)
$$

and therefore

$$
\tilde{X}_{2111}^{\prime}=-\frac{1}{480}\left(\delta_{m}^{[b}\left(u_{1} \gamma^{c]} \gamma^{d e} u_{2}\right)+\delta_{m}^{[d}\left(u_{1} \gamma^{e]} \gamma^{b c} u_{2}\right)\right)
$$

The result for $\tilde{X}_{4010}$ is

$$
\begin{aligned}
\tilde{X}_{4010}=[- & \frac{1}{360} \delta_{m n}^{b q}\left(u_{1} \gamma^{c} u_{2}\right)-\frac{1}{90} \delta_{m q}^{b c}\left(u_{1} \gamma^{n} u_{2}\right)+\frac{1}{720} \delta_{m n}^{b c}\left(u_{1} \gamma^{q} u_{2}\right)-\frac{1}{2520} \delta_{q}^{m}\left(u_{1} \gamma^{b c n} u_{2}\right) \\
& \left.-\frac{1}{720} \delta_{q}^{b}\left(u_{1} \gamma^{c m n} u_{2}\right)+\frac{1}{1260} \delta_{m}^{b}\left(u_{1} \gamma^{c n q} u_{2}\right)+\frac{1}{3360}\left(u_{1} \gamma^{b c m n q} u_{2}\right)\right]_{[b c][m n]}
\end{aligned}
$$

For the evaluation of $\tilde{X}_{2210}$, it is useful to consider the more general correlator

$$
\begin{aligned}
&\left\langle\left(\lambda \gamma^{a} \theta\right)\left(\theta \gamma_{a} u_{1}\right)\left(\lambda \gamma^{[m \mid} \gamma^{b c} \theta\right)\left(\lambda \gamma^{\mid n]} \gamma^{d e} \theta\right)\left(\theta \gamma_{x} u_{2}\right)\right\rangle=\left[-\frac{13}{5040} \delta_{x}^{d} \delta_{m n}^{b e}\left(u_{1} \gamma^{c} u_{2}\right)+\ldots\right. \\
&\left.+\frac{11}{201600} \delta_{x}^{m}\left(u_{1} \gamma^{b c d e n} u_{2}\right)+\cdots-\frac{11}{403200}\left(u_{1} \gamma^{b c d e m n x} u_{2}\right)\right]_{[m n][b c][d e]} \\
& \quad+\frac{1}{9676800} \varepsilon_{b c d e m n i_{1} i_{2} i_{3} i_{4}}\left(u_{1} \gamma^{i_{1} i_{2} i_{3} i_{4} x} u_{2}\right)-\frac{1}{2419200} \varepsilon_{b c d e m n x i_{1} i_{2} i_{3}}\left(u_{1} \gamma^{i_{1} i_{2} i_{3}} u_{2}\right)
\end{aligned}
$$

This time, even using the method of section 2.2, there are sufficiently many open indices and long enough traces for epsilon tensors to appear. Using eqs. (2.11) and (2.12), they can be re-written into $\gamma^{[5,7]}$ terms:

$$
\begin{gather*}
\left\langle\left(\lambda \gamma^{a} \theta\right)\left(\theta \gamma_{a} u_{1}\right)\left(\lambda \gamma^{[m \mid} \gamma^{b c} \theta\right)\left(\lambda \gamma^{\mid n]} \gamma^{d e} \theta\right)\left(\theta \gamma_{x} u_{2}\right)\right\rangle=\left[-\frac{13}{5040} \delta_{x}^{d} \delta_{m n}^{b e}\left(u_{1} \gamma^{c} u_{2}\right)+\ldots\right. \\
\left.+\frac{1}{16800} \delta_{x}^{m}\left(u_{1} \gamma^{b c d e n} u_{2}\right)+\cdots-\frac{1}{33600}\left(u_{1} \gamma^{b c d e m n x} u_{2}\right)\right]_{[m n][b c][d e]} \tag{27terms}
\end{gather*}
$$

A good check on the sign of the epsilon contributions is that $\tilde{X}_{2111}^{\prime}$ is recovered when contracting with $\eta_{n x}$, involving a cancellation of all $\gamma^{[5]}$ terms. To obtain $\tilde{X}_{2210}$, one multiplies by $-\eta_{c x}$ :

$$
\begin{aligned}
\tilde{X}_{2210}=[ & \frac{1}{720} \delta_{m n}^{d e}\left(u_{1} \gamma^{b} u_{2}\right)+\frac{29}{2880} \delta_{m n}^{b d}\left(u_{1} \gamma^{e} u_{2}\right)+\frac{11}{2880} \delta_{d e}^{b m}\left(u_{1} \gamma^{n} u_{2}\right)+\frac{1}{20160} \delta_{m}^{d}\left(u_{1} \gamma^{b e n} u_{2}\right) \\
& \left.+\frac{1}{2880} \delta_{m}^{b}\left(u_{1} \gamma^{d e n} u_{2}\right)+\frac{11}{20160} \delta_{d}^{b}\left(u_{1} \gamma^{e m n} u_{2}\right)+\frac{1}{4480}\left(u_{1} \gamma^{b d e m n} u_{2}\right)\right]_{[d e][m n]}
\end{aligned}
$$

For the calculation of $X_{2030}$ and $X_{2012}$, one may first evaluate a more general correlator $\left\langle\left(\lambda \gamma^{a} \theta\right)\left(\theta \gamma_{a} u_{1}\right)\left(\lambda \gamma^{[m} u_{2}\right)\left(\lambda \gamma^{n]} \gamma^{b c} \theta\right)\left(\theta \gamma^{x} \gamma^{d e} \theta\right)\right\rangle$ and then contract with $\eta_{c x}$ and $\eta_{n x}$, respectively. The results are

$$
\begin{aligned}
\tilde{X}_{2030}= & {\left[-\frac{1}{720} \delta_{m n}^{d e}\left(u_{1} \gamma^{b} u_{2}\right)+\frac{1}{288} \delta_{m n}^{b d}\left(u_{1} \gamma^{e} u_{2}\right)-\frac{1}{1440} \delta_{d e}^{b m}\left(u_{1} \gamma^{n} u_{2}\right)-\frac{17}{10080} \delta_{m}^{d}\left(u_{1} \gamma^{b e n} u_{2}\right)\right.} \\
& \left.-\frac{23}{10080} \delta_{m}^{b}\left(u_{1} \gamma^{d e n} u_{2}\right)-\frac{1}{1440} \delta_{d}^{b}\left(u_{1} \gamma^{e m n} u_{2}\right)+\frac{1}{6720}\left(u_{1} \gamma^{b d e m n} u_{2}\right)\right]_{[m n][d e]} \\
\tilde{X}_{2012}= & {\left[\frac{1}{288} \delta_{b m}^{d e}\left(u_{1} \gamma^{c} u_{2}\right)+\frac{1}{288} \delta_{d m}^{b c}\left(u_{1} \gamma^{e} u_{2}\right)-\frac{1}{1440} \delta_{d e}^{b c}\left(u_{1} \gamma^{m} u_{2}\right)+\frac{1}{2016} \delta_{m}^{d}\left(u_{1} \gamma^{b c e} u_{2}\right)\right.} \\
& \left.-\frac{11}{10080} \delta_{m}^{b}\left(u_{1} \gamma^{c d e} u_{2}\right)+\frac{17}{10080} \delta_{d}^{b}\left(u_{1} \gamma^{c e m} u_{2}\right)-\frac{1}{3360}\left(u_{1} \gamma^{b c d e m} u_{2}\right)\right]_{[b c][d e]}
\end{aligned}
$$

After multiplication with the momenta and polarisations, all individual contributions are gauge invariant and can be expanded in the basis $C_{1} \ldots C_{5}$ listed in (A.6):

$$
\begin{aligned}
\left(1-\pi_{34}\right) X_{4010} & =\frac{1}{60480}(-1,12,-12,1,2)_{C_{1} \ldots C_{5}} \\
\left(1-\pi_{34}\right) X_{2210} & =\frac{1}{60480}(31,-30,30,-1,-2)_{C_{1} \ldots C_{5}} \\
\left(1-\pi_{34}\right) X_{2030} & =\frac{1}{60480}(0,-21,21,0,0)_{C_{1} \ldots C_{5}} \\
\left(1-\pi_{34}\right) X_{2012} & =\frac{1}{60480}(-30,39,-39,0,0)_{C_{1} \ldots C_{5}} \\
X_{2111}^{\prime} & =\frac{1}{480}(1,0,4,-1,0)_{C_{1} \ldots C_{5}}
\end{aligned}
$$

As expected, the terms with the wrong [34] symmetry cancel:

$$
\left(1-\pi_{34}\right)\left(X_{4010}+X_{2210}+X_{2030}+X_{2012}\right)=0
$$

The remaining part can be written as

$$
\begin{equation*}
K_{1-\mathrm{loop}}^{2 \mathrm{~B} 2 \mathrm{~F}}=X_{2111}^{\prime}=-\frac{1}{480}\left(s_{13}\left(u_{2} \not ぬ_{3}\left(\not k_{2}+\not k_{3}\right) \not \not_{4} u_{1}\right)+s_{23}\left(u_{2} \not_{4}\left(\not k_{1}+\not k_{3}\right) \not_{3} u_{1}\right)\right) \tag{3.6}
\end{equation*}
$$

This again agrees with amplitude computed in the RNS result, see [16] eq. (3.37).

## 4. Two-loop amplitudes

The pure spinor formalism was used in [4, 2] to compute the two-loop type-IIB amplitude involving four massless states,

$$
\mathcal{A}=\int d^{2} \Omega_{11} d^{2} \Omega_{12} d^{2} \Omega_{22} \prod_{i=1}^{4} \int d^{2} z_{i} \frac{\exp \left(-\sum_{i, j} k_{i} \cdot k_{j} G\left(z_{i}, z_{j}\right)\right)}{(\operatorname{det} \operatorname{Im} \Omega)^{5}} K_{2-\operatorname{loop}}\left(k_{i}, z_{i}\right)
$$

where $\Omega$ is the genus-two period matrix, and the integration over fermionic zero modes is encapsulated in

$$
\begin{align*}
K_{2 \text {-loop }} & =\Delta_{12} \Delta_{34}\left\langle\left(\lambda \gamma^{m n p q r} \lambda\right)\left(\lambda \gamma^{s} W_{1}\right) \mathcal{F}_{2, m n} \mathcal{F}_{3, p q} \mathcal{F}_{4, r s}\right\rangle+(\operatorname{perm}(1234))  \tag{4.1}\\
& \equiv \Delta_{12} \Delta_{34} K_{12}+\Delta_{13} \Delta_{24} K_{13}+\Delta_{14} \Delta_{23} K_{14} . \tag{4.2}
\end{align*}
$$

The kinematic factors $K_{12}, K_{13}, K_{14}$ are accompanied by the basic antisymmetric biholomorphic 1-form $\Delta$, which is related to a canonical basis $\omega_{1}, \omega_{2}$ of holomorphic differentials via $\Delta_{i j}=\Delta\left(z_{i}, z_{j}\right)=\omega_{1}\left(z_{i}\right) \omega_{2}\left(z_{j}\right)-\omega_{2}\left(z_{i}\right) \omega_{1}\left(z_{j}\right)$. The superfields $W_{i}^{\alpha}$ and $\mathcal{F}_{i, m n}$ are the spinor and vector field strengths of the $i$-th external state, as in section 3. One encounters superspace integrals of the form

$$
\begin{equation*}
Y(a b c d)=\left\langle\left(\lambda \gamma^{m n p q r} \lambda\right)\left(\lambda \gamma^{s} W_{a}\right) \mathcal{F}_{b, m n} \mathcal{F}_{c, p q} \mathcal{F}_{d, r s}\right\rangle \tag{4.3}
\end{equation*}
$$

The symmetries of the $\lambda^{3}$ combination [4] in this correlator include the obvious symmetry under $m n \leftrightarrow p q$, and also $\left(\lambda \gamma^{[m n p q r} \lambda\right)\left(\lambda \gamma^{s]}\right)_{\alpha}=0$ (this holds for pure spinors $\lambda$ and can be seen by dualising, and holds for unconstrained spinors $\lambda$ as part of a $\lambda^{3} \theta^{5}$ scalar, as seen from the representation content (2.8)), and allow one to shuffle the $\mathcal{F}$ factors:

$$
\begin{equation*}
Y(a b c d)=Y(a c b d), \quad Y(a b c d)+Y(a c d b)+Y(a d b c)=0 . \tag{4.4}
\end{equation*}
$$

This applies also after expanding the superfields in powers of $\theta$ : Using the notation

$$
Y_{A B C D}(a b c d)=\left\langle\left(\lambda \gamma^{m n p q r} \lambda\right)\left(\lambda \gamma^{s} W_{a}^{(A)}\right) \mathcal{F}_{b, m n}^{(B)} \mathcal{F}_{c, p q}^{(C)} \mathcal{F}_{d, r s}^{(D)}\right\rangle,
$$

one obtains the identities

$$
\begin{align*}
& Y_{A B C D}(a b c d)=(-)^{B C} Y_{A C B D}(a c b d) \\
& Y_{A B C D}(a b c d)=(-)^{B C+B D+1} Y_{A C D B}(a c d b)+(-)^{B D+C D+1} Y_{A D B C}(a d b c), \tag{4.5}
\end{align*}
$$

where the signs arise from the exchange of fermions.

### 4.1 Review: four bosons

The case of four Neveu-Schwarz states was considered in [6] and will be briefly reviewed here. As all three kinematic factors $K_{12}, K_{13}$ and $K_{14}$ are equivalent, it is sufficient to consider $K_{12}$ in detail. With all external states being identical, the symmetrisations of (4.1) can be carried out at the end of the calculation:

$$
\begin{aligned}
K_{12}^{4 \mathrm{~B}} & =-\left.4\left\langle W_{[1} \mathcal{F}_{2]} \mathcal{F}_{[3} \mathcal{F}_{4]}\right\rangle\right|_{4 \mathrm{~B}}-\left.4\left\langle W_{[3} \mathcal{F}_{4]} \mathcal{F}_{[1} \mathcal{F}_{2]}\right\rangle\right|_{4 \mathrm{~B}} \\
& =\left.\left(1-\pi_{12}\right)\left(1-\pi_{34}\right)\left(1+\pi_{13} \pi_{24}\right)\left\langle W_{1} \mathcal{F}_{2} \mathcal{F}_{3} \mathcal{F}_{4}\right\rangle\right|_{4 \mathrm{~B}}
\end{aligned}
$$

Expanding the superfields, the Neveu-Schwarz states come from terms of the form $Y_{A B C D} \equiv$ $Y_{A B C D}(1234)$ with $A$ odd and $B, C, D$ even. Using the shuffling identities (4.5) to simplify, one obtains

$$
\begin{aligned}
\left.\left\langle W_{1} \mathcal{F}_{2} \mathcal{F}_{3} \mathcal{F}_{4}\right\rangle\right|_{4 \mathrm{~B}} & =Y_{5000}+Y_{1400}+Y_{1040}+Y_{1004}+Y_{3200}+Y_{3020}+Y_{3002}+Y_{1220}+Y_{1202}+Y_{1022} \\
& =\left(1+\pi_{23}\right)\left(1-\pi_{24}\right)\left(\frac{1}{3} Y_{5000}+Y_{1400}+Y_{3200}+Y_{1022}\right)
\end{aligned}
$$

and therefore $K_{12}^{4 \mathrm{~B}}$ can be written as the image of a symmetrisation operator $\mathcal{S}_{4 \mathrm{~B}}$

$$
\begin{aligned}
K_{12}^{4 \mathrm{~B}} & =\mathcal{S}_{4 \mathrm{~B}}\left(\frac{1}{3} Y_{5000}+Y_{1400}+Y_{3200}+Y_{1022}\right) \\
\mathcal{S}_{4 \mathrm{~B}} & =\left(1-\pi_{12}\right)\left(1-\pi_{34}\right)\left(1+\pi_{13} \pi_{24}\right)\left(1+\pi_{23}\right)\left(1-\pi_{24}\right)
\end{aligned}
$$

It is worth noting at this point that, on the sixteen-dimensional space of Lorentz scalars built from the four field strengths $F_{i}$ and two momenta, the symmetriser $\mathcal{S}_{4 \mathrm{~B}}$ has rank four. The correlators were computed in [6], using the method outlined in section 2.1. Two are zero, $Y_{5000}=Y_{1400}=0$, and the remaining ones are

$$
\begin{aligned}
Y_{3200} & =-\frac{1}{192} k_{a}^{1} F_{c d}^{1} k_{m}^{2} F_{e f}^{2} F_{p q}^{3} F_{r s}^{4}\left\langle\left(\lambda \gamma^{m n p q r} \lambda\right)\left(\lambda \gamma^{s} \gamma^{a b} \theta\right)\left(\theta \gamma_{b}^{c d} \theta\right)\left(\theta \gamma_{n}^{e f} \theta\right)\right\rangle \\
Y_{1022} & =-\frac{1}{64} F_{a b}^{1} F_{m n}^{2} k_{p}^{3} F_{c d}^{3} k_{r}^{4} F_{e f}^{4}\left\langle\left(\lambda \gamma^{m n p q[r} \lambda\right)\left(\lambda \gamma^{s]} \gamma^{a b} \theta\right)\left(\theta \gamma_{q}^{c d} \theta\right)\left(\theta \gamma_{s}^{e f} \theta\right)\right\rangle
\end{aligned}
$$

In reducing those two contributions to a set of independent scalars, one finds that they both are not just sums of $(k \cdot k) F^{4}$ terms but also contain terms of the form $k \cdot F$ terms. The latter are projected out by the symmetriser $\mathcal{S}_{4 \mathrm{~B}}$, and the result is

$$
\begin{aligned}
\mathcal{S}_{4 \mathrm{~B}}\left(Y_{3200}+Y_{1022}\right) & =-\frac{1}{120}\left(s_{13}-s_{23}\right)\left(4 \operatorname{Tr}\left(F_{(1} F_{2} F_{3} F_{4)}\right)-\operatorname{Tr}\left(F_{(1} F_{2}\right) \operatorname{Tr}\left(F_{3} F_{4)}\right)\right) \\
K_{12}^{4 \mathrm{~B}} & =-\frac{1}{720}\left(s_{13}-s_{23}\right) t_{8} F^{4}
\end{aligned}
$$

By trivial index exchange, one obtains $K_{13}$ and $K_{14}$, and the total is

$$
\begin{equation*}
K_{2-\text { loop }}^{4 \mathrm{~B}}=-\frac{1}{720}\left(\left(s_{13}-s_{23}\right) \Delta_{12} \Delta_{34}+\left(s_{12}-s_{23}\right) \Delta_{13} \Delta_{24}+\left(s_{12}-s_{13}\right) \Delta_{14} \Delta_{23}\right) t_{8} F^{4} \tag{4.6}
\end{equation*}
$$

a product of the completely symmetric one-loop kinematic factor $t_{8} F^{4}$ and a completely symmetric combination of the momenta and the $\Delta_{i j}$.

### 4.2 Four fermions

The calculation involving four Ramond states is very similar to the bosonic one. Focussing on the $K_{12}$ part, the symmetrisations in (4.1) can again be rewritten as action of symmetrisation operators on the correlator of superfields with one particular labelling:

$$
\begin{aligned}
K_{12}^{4 \mathrm{~F}} & =\left.\left(1-\pi_{12}\right)\left(1-\pi_{34}\right)\left(1+\pi_{13} \pi_{24}\right)\left\langle W_{1} \mathcal{F}_{2} \mathcal{F}_{3} \mathcal{F}_{4}\right\rangle\right|_{4 \mathrm{~F}} \\
& =\left.4\left(1-\pi_{12}\right)\left\langle W_{1} \mathcal{F}_{2} \mathcal{F}_{3} \mathcal{F}_{4}\right\rangle\right|_{4 \mathrm{~F}}
\end{aligned}
$$

The last step follows from the fact that all $k^{4} u^{4}$ scalars (see appendix A.2) are invariant under $\pi_{13} \pi_{24}$ and have $\pi_{12}=\pi_{34}$. This time, on expanding the superfields, one collects the terms $Y_{A B C D}$ with $A$ even and $B, C, D$ odd. After using (4.4) to simplify,

$$
\begin{aligned}
\left.\left\langle W_{1} \mathcal{F}_{2} \mathcal{F}_{3} \mathcal{F}_{4}\right\rangle\right|_{4 \mathrm{~F}} & =Y_{2111}+Y_{0311}+Y_{0131}+Y_{0113} \\
& =\left(1-\pi_{23}\right)\left(1+\pi_{24}\right)\left(\frac{1}{3} Y_{2111}+Y_{0311}\right),
\end{aligned}
$$

one obtains

$$
K_{12}^{4 \mathrm{~F}}=\mathcal{S}_{4 \mathrm{~F}}\left(\frac{1}{3} Y_{2111}+Y_{0311}\right), \quad \mathcal{S}_{4 \mathrm{~F}}=4\left(1-\pi_{12}\right)\left(1-\pi_{23}\right)\left(1+\pi_{24}\right)
$$

This symmetriser has rank three, and the result is again not determined by symmetry. Two correlators have to be computed:

$$
\begin{aligned}
& Y_{2111}=(-2) k_{a}^{1} k_{m}^{2} k_{p}^{3} k_{r}^{4}\left\langle\left(\lambda \gamma^{m n p q[r} \lambda\right)\left(\lambda \gamma^{s]} \gamma^{a b} \theta\right)\left(\theta \gamma_{b} u_{1}\right)\left(\theta \gamma_{n} u_{2}\right)\left(\theta \gamma_{q} u_{3}\right)\left(\theta \gamma_{s} u_{4}\right)\right\rangle \\
& Y_{0311}=\left(-\frac{2}{3}\right) k_{a}^{2} k_{m}^{2} k_{p}^{3} k_{r}^{4}\left\langle\left(\lambda \gamma^{m n p q[r} \lambda\right)\left(\lambda \gamma^{s} u_{1}\right)\left(\theta \gamma_{n}^{a b} \theta\right)\left(\theta \gamma_{b} u_{2}\right)\left(\theta \gamma_{q} u_{3}\right)\left(\theta \gamma_{s} u_{4}\right)\right\rangle
\end{aligned}
$$

With four fermions present, the method of section 2.2 is preferred as it does not involve rearranging the fermions using Fierz identities. The first correlator was covered as an example in that section, and the second one can be evaluated in the same fashion. Expressed in the basis listed in (A.5), the results are

$$
\begin{aligned}
& Y_{2111}=\frac{1}{5040}(-19,-21,21,19,-17,-17,0,0,0,0)_{B_{1} \ldots B_{10}} \\
& Y_{0311}=\frac{1}{15120}(-14,-16,0,2,-8,-8,0,-5,-5,0)_{B_{1} \ldots B_{10}} .
\end{aligned}
$$

After acting with the symmetriser $\mathcal{S}_{4 \mathrm{~F}}$, one obtains the same $u^{4}$ scalar encountered in the one-loop amplitude,

$$
\begin{aligned}
K_{12}^{4 \mathrm{~F}} & =\mathcal{S}_{4 \mathrm{~F}}\left(\frac{1}{3} Y_{2111}+Y_{0311}\right)=\frac{1}{45}(-1,0,1,0,-1,0,0,0,0,0)_{B_{1} \ldots B_{10}} \\
& =\frac{1}{45} s_{12}\left(\left(u_{1} \not k_{3} u_{2}\right)\left(u_{3} k_{1} u_{4}\right)-\left(u_{1} \not k_{2} u_{3}\right)\left(u_{2} \not k_{1} u_{4}\right)+\left(u_{1} \not k_{2} u_{4}\right)\left(u_{2} \not k_{1} u_{3}\right)\right) .
\end{aligned}
$$

The $K_{13}$ and $K_{14}$ parts again follow by index exchange, and the total result

$$
\begin{align*}
K_{2-\mathrm{loop}}^{4 \mathrm{~F}}=\frac{1}{45} \cdot\left(s_{12} \Delta_{12} \Delta_{34}-s_{13} \Delta_{13} \Delta_{24}+\right. & \left.s_{14} \Delta_{14} \Delta_{23}\right) \\
& \times\left(\left(u_{1} \not k_{3} u_{2}\right)\left(u_{3} k_{1} u_{4}\right)-\left(u_{1} \not k_{2} u_{3}\right)\left(u_{2} \not k_{1} u_{4}\right)+\left(u_{1} \not k_{2} u_{4}\right)\left(u_{2} \not k_{1} u_{3}\right)\right) \tag{4.7}
\end{align*}
$$

is again a simple product of the one-loop kinematic factor and a combination of the $\Delta_{i j}$ and momenta. This time, both factors are completely antisymmetric.

### 4.3 Two bosons, two fermions

As in the one-loop calculation of section 3.3, in the mixed case one has to pay some attention to the permutations in (4.1) since they affect which superfields contribute fermionic fields. The complete symmetrisation makes it irrelevant which labels are assigned to the two fermions, and the convention $f_{1} f_{2} b_{3} b_{4}$ will be used here. The kinematic factor $K_{12}^{2 \mathrm{~B} 2 \mathrm{~F}}$ is then distinguished from the other two, $K_{13}^{2 \mathrm{~B} 2 \mathrm{~F}}$ and $K_{14}^{2 \mathrm{~B} 2 \mathrm{~F}}$.

Some elementary symmetry considerations prevent the $K_{12}^{2 \mathrm{~B} 2 \mathrm{~F}} \ldots K_{14}^{2 \mathrm{~B} 2 \mathrm{~F}}$ to have the same simple form as in the other two cases, i.e. a linear function of the Mandelstam invariants, times the one-loop kinematic factor:

$$
K_{2 \text {-loop }}^{2 \mathrm{~B} 2 \mathrm{~F}} \neq \sum_{j=1,2,3} \Delta_{1 j} \Delta_{k l} \times\left[f_{j}\left(s_{12}, s_{13}\right) K_{1-\text { loop }}^{2 \mathrm{~B} 2 \mathrm{~F}}\left(f_{1} f_{2} b_{3} b_{4}\right)\right] .
$$

For $K_{12}$, this is the case because the coefficient of $\Delta_{12} \Delta_{34}$ in (4.2) is antisymmetric both under $1 \leftrightarrow 2$ and $3 \leftrightarrow 4$. On the other hand, the one-loop kinematic factor $K_{1 \text {-loop }}^{4 \mathrm{~F}}\left(f_{1} f_{2} b_{3} b_{4}\right)$ has the symmetries [12] and (34), and any function $f\left(s_{12}, s_{13}\right)$ has identical symmetry under $1 \leftrightarrow 2$ and $3 \leftrightarrow 4$. Similarly, in the case of $K_{13}$, note that any $f\left(s_{12}, s_{13}\right)$ is symmetric under the simultaneous interchange $1 \leftrightarrow 2$ and $3 \leftrightarrow 4$, whereas $K_{1-\text { loop }}^{2 \mathrm{~B} 2 \mathrm{~F}}\left(f_{1} f_{2} b_{3} b_{4}\right)$ is antisymmetric.

Carrying out the symmetrisations in (4.1) and using the identities (4.4), one finds

$$
\begin{aligned}
K_{12}^{2 \mathrm{~B} 2 \mathrm{~F}} & =\left(\mathbf{1}-\pi_{12}\right)\left(\mathbf{1}-\pi_{34}\right) \tilde{K}, \\
K_{13}^{2 \mathrm{~B} 2 \mathrm{~F}} & =\left(2 \cdot \mathbf{1}+\pi_{12}+\pi_{34}+2 \pi_{12} \pi_{34}\right) \tilde{K}, \\
K_{14}^{2 \mathrm{~B} 2 \mathrm{~F}} & =\left(\mathbf{1}+2 \pi_{12}+2 \pi_{34}+\pi_{12} \pi_{34}\right) \tilde{K},
\end{aligned}
$$

where, schematically,

$$
\begin{equation*}
\tilde{K}=\left\langle\left(\lambda^{3} W_{1}^{(\text {even })}\right) \mathcal{F}_{2}^{(\text {odd })} \mathcal{F}_{3}^{(\text {even })} \mathcal{F}_{4}^{(\text {even })}\right\rangle+\left\langle\left(\lambda^{3} W_{3}^{(\text {odd })}\right) \mathcal{F}_{4}^{(\text {even })} \mathcal{F}_{1}^{(\text {odd })} \mathcal{F}_{2}^{(\text {odd })}\right\rangle \tag{4.8}
\end{equation*}
$$

Considering the action of these symmetrisation operators on the ten-dimensional space of $k^{3} u_{1} u_{2} F_{3} F_{4}$ scalars (see appendix A.3), one finds that $\left(\mathbf{1}-\pi_{12}\right)\left(\mathbf{1}-\pi_{34}\right)$ has rank one with image spanned by $s_{12} \times(1,-2,2,0,0)_{C_{1} \ldots C_{5}}$, and both $\left(2 \cdot \mathbf{1}+\pi_{12}+\pi_{34}+2 \pi_{12} \pi_{34}\right)$ and $\left(\mathbf{1}+2 \pi_{12}+2 \pi_{34}+\pi_{12} \pi_{34}\right)$ have rank 2 with images spanned by $\left(s_{12}\right.$ and $\left.s_{13}\right) \times$ $(1,-2,2,0,0)_{C_{1} \ldots C_{5}}$. Here,

$$
(1,-2,2,0,0)_{C_{1} \ldots C_{5}}=\left(u_{1} \gamma^{a} u_{2}\right)\left(k_{a}^{3} F_{b c}^{3} F_{b c}^{4}-2 F_{a b}^{3} F_{b c}^{4} k_{c}^{1}+2 F_{a b}^{4} F_{b c}^{3} k_{c}^{1}\right)
$$

is the unique (up to normalisation) scalar of the form $k^{3} u_{1} u_{2} F_{3} F_{4}$ antisymmetric in both [12] and [34]. The tensor structure is thus fixed by elementary symmetries to be

$$
\begin{align*}
K_{2 \text {-loop }}^{2 \mathrm{~B} 2 \mathrm{~F}}=\left(c_{1} s_{12} \Delta_{12} \Delta_{34}+\left(c_{2} s_{12}\right.\right. & \left.\left.+c_{3} s_{13}\right) \Delta_{13} \Delta_{24}+\left(\left(c_{3}-c_{2}\right) s_{12}+c_{3} s_{13}\right) \Delta_{14} \Delta_{23}\right) \\
& \times\left(u_{1} \gamma^{a} u_{2}\right)\left(k_{a}^{3} F_{b c}^{3} F_{b c}^{4}-2 F_{a b}^{3} F_{b c}^{4} k_{c}^{1}+2 F_{a b}^{4} F_{b c}^{3} k_{c}^{1}\right) \tag{4.9}
\end{align*}
$$

where the $c_{i}$ are constants which could be determined by calculating the non-zero correlators in (4.8).

## 5. Discussion

In this paper, different methods were discussed to efficiently evaluate the superspace integrals appearing in multiloop amplitudes derived in the pure spinor formalism. Extending previous calculations [6, 7] restricted to Neveu-Schwarz states, it was then shown how the treatment of Ramond states poses no additional difficulties. It was seen that simple exchange symmetries completely determine the kinematic factor of the one-loop four-fermion amplitude, and are sufficient to derive the two-boson/two-fermion two-loop kinematic factor starting from symmetry properties of the corresponding pure spinor integral.

While the bosonic calculations of [6, 7] have, in conjunction with supersymmetry, already established the equivalence of the massless four-point amplitudes derived in the pure spinor and RNS formalisms, it would be interesting to make contact between the results of sections 4.2 / 4.3 and two-loop amplitudes involving Ramond states as computed in the RNS formalism (see for example [18]).

The assistance of a computer algebra system seems indispensible in explicitly evaluating pure spinor superspace integrals. To avoid excessive use of custom-made algorithms, it would be desirable to implement these calculations in a wider computational framework particular adapted to field theory calculations (19].

The methods outlined in this paper should be easily applicable to future higher-loop amplitude expressions derived from the pure spinor formalism, and, it is hoped, to other superspace integrals.

## Acknowledgments

The author would like to thank Louise Dolan for discussions. This work is supported by the U.S. Department of Energy, grant no. DE-FG01-06ER06-01, Task A.

## A. Reduction to kinematic bases

In calculating scattering amplitudes one encounters kinematic factors which are Lorentz invariant polynomials in the momenta, polarisations and/or spinor wavefunctions of the scattered particles. It can be a non-trivial task to simplify such expressions, taking into account the on-shell identities $\sum_{i} k_{i}=0, k_{i}^{2}=0, k_{i} \cdot \zeta_{i}=0, \not k_{i} u_{i}=0$, and, in the case of fermions, re-arrangements stemming from Fierz identities.

More generally, one would like to know how many independent combinations of some given fields (subject to on-shell identities) there are, and how to reduce an arbitrary expression with respect to some chosen basis. This appendix outlines methods to address these problems, with an emphasis on algorithms which can easily be transferred to a computer algebra system. These methods are not limited to dealing with pure spinor calculations but the scope will be restricted to amplitudes of four massless vector or spinor particles in ten dimensions.

## A. 1 Four bosons

It is not difficult to reduce polynomials in the momenta and polarisations to a canonical form. The momentum conservation constraint $\sum_{i} k_{i}=0$ is solved by eliminating one momentum (for example $k_{4}$ ), all $k_{i}^{2}$ are set to zero, and one of the two remaining quadratic combinations of momenta is eliminated (for example $s_{23} \rightarrow-s_{12}-s_{13}$, where $s_{i j} \equiv k_{i} \cdot k_{k}$ ). Then all products $k_{i} \cdot \zeta_{i}$ are set to zero, and one extra $k \cdot \zeta$ product is replaced (when eliminating $k_{4}$, the replacement is $\left.k_{3} \cdot \zeta_{4} \rightarrow\left(-k_{1}-k_{2}\right) \cdot \zeta_{4}\right)$. The remaining monomials are then independent. (This is at least the case with the low powers of momenta encountered in the calculations of sections 3 and 4, where there are enough spatial directions for all momenta/polarisations to be linearly independent.)

The implementation of these reduction rules on a computer is straightforward. The easiest way to obtain scalars which are also invariant under the gauge symmetry $k_{i} \rightarrow \zeta_{i}$ is to start with expressions constructed from the field strengths $F_{i}^{a b}=2 k_{i}^{[a} \zeta_{i}^{b]}$. For the one-loop calculations of section 3.1, the relevant basis consists of gauge invariant scalars containing only the four field strengths $F_{1} \ldots F_{4}$. One finds six independent combinations,

$$
\begin{array}{ll}
\operatorname{Tr}\left(F_{1} F_{2} F_{3} F_{4}\right) & \operatorname{Tr}\left(F_{1} F_{2}\right) \operatorname{Tr}\left(F_{3} F_{4}\right) \\
\operatorname{Tr}\left(F_{1} F_{2} F_{4} F_{3}\right) & \operatorname{Tr}\left(F_{1} F_{3}\right) \operatorname{Tr}\left(F_{2} F_{4}\right) \\
\operatorname{Tr}\left(F_{1} F_{3} F_{2} F_{4}\right) & \operatorname{Tr}\left(F_{1} F_{4}\right) \operatorname{Tr}\left(F_{2} F_{3}\right)
\end{array}
$$

In the two-loop calculations of section 4.1, all monomials have two more momenta. There are sixteen independent gauge invariant scalars of the form $k k F_{1} F_{2} F_{3} F_{4}$, and twelve of them may be constructed from the previous basis by multiplication with $s_{12}$ and $s_{13}$ : $A_{1}=s_{12} \operatorname{Tr}\left(F_{1} F_{2} F_{3} F_{4}\right), A_{2}=s_{13} \operatorname{Tr}\left(F_{1} F_{2} F_{3} F_{4}\right)$, etc. One choice for the additional four is

$$
\begin{array}{ll}
A_{13}=k_{3} \cdot F_{1} \cdot F_{2} \cdot k_{3} \operatorname{Tr}\left(F_{3} F_{4}\right) & A_{15}=k_{3} \cdot F_{1} \cdot F_{4} \cdot k_{2} \operatorname{Tr}\left(F_{2} F_{3}\right) \\
A_{14}=k_{4} \cdot F_{1} \cdot F_{3} \cdot k_{2} \operatorname{Tr}\left(F_{2} F_{4}\right) & A_{16}=k_{4} \cdot F_{2} \cdot F_{3} \cdot k_{4} \operatorname{Tr}\left(F_{1} F_{4}\right) .
\end{array}
$$

As an example application of the computer algorithms, one may check that the symmetrisation operator of section 4.1,

$$
\mathcal{S}_{4 \mathrm{~B}}=\left(1-\pi_{12}\right)\left(1-\pi_{34}\right)\left(1+\pi_{13} \pi_{24}\right)\left(1+\pi_{23}\right)\left(1-\pi_{24}\right)
$$

acts as

$$
\begin{aligned}
\mathcal{S}_{4 \mathrm{~B}} A_{1} & =8 A_{1}+4 A_{2}-4 A_{3}+4 A_{4}+8 A_{5}+16 A_{6} \\
& \cdots \\
\mathcal{S}_{4 \mathrm{~B}} A_{16} & =-6 A_{1}+6 A_{3}-6 A_{5}-12 A_{6}+\frac{3}{2} A_{7}+3 A_{8}+\frac{3}{2} A_{9}+3 A_{10}+\frac{3}{2} A_{11}+3 A_{12}
\end{aligned}
$$

and has rank four.

## A. 2 Four fermions

In dealing with the spinor wavefunctions $u_{i}$ one has to face two issues: Fierz identities, and the Dirac equation. Fierz identities not only allow one to change the order of the spinors
but also give rise to relations between different expressions in one spinor order. The Dirac equation often simplifies terms with momenta contracted into $\left(u_{i} \gamma^{[n]} u_{j}\right)$ bilinears.

In this section it is shown how to construct bases for terms of the form $\left(k^{2}\right.$ or $\left.k^{4}\right) \times$ $u_{1} u_{2} u_{3} u_{4}$. A significant simplification comes from noting that the Dirac equation allows one to rewrite ( $u_{i} \gamma^{[n]} u_{j}$ ) bilinears into terms with lower $n$ if more than one momentum is contracted into the $\gamma^{[n]}$. A good first step is therefore to disregard the momenta temporarily and find all independent scalars and two-index tensors built from $u_{1}, \ldots, u_{4}$. From the $\mathrm{SO}(10)$ representation content,

$$
\left(S^{+}\right)^{\otimes 4}=2 \cdot \mathbf{1}+6 \cdot \text { 日 }+3 \cdot \widetilde{\square}+(\text { tensors with rank }>2),
$$

one expects two scalars and nine 2 -tensors. The scalars are easily found by considering, as in (20],

$$
\begin{aligned}
& T_{1}(1234)=\left(u_{1} \gamma^{a} u_{2}\right)\left(u_{3} \gamma_{a} u_{4}\right), \\
& T_{3}(1234)=\left(u_{1} \gamma^{a b c} u_{2}\right)\left(u_{3} \gamma_{a b c} u_{4}\right) .
\end{aligned}
$$

and similarly for the other two inequivalent orders of the four spinors. (Note there is no $T_{5}$ because of self-duality of the $\gamma^{[5]}$.) From Fierz transformations, one learns that all $T_{3}$ terms can be reduced to $T_{1}$ by $T_{3}(1234)=-12 T_{1}(1234)-24 T_{1}(1324)$ and permutations, and the identity $\left(\gamma_{a}\right)_{(\alpha \beta}\left(\gamma^{a}\right)_{\gamma) \delta}=0$ implies that $T_{1}(1234)+T_{1}(1324)+T_{1}(1423)=0$, leaving for example $T_{1}$ (1234) and $T_{1}$ (1324) as independent scalars.

Generalising this approach to two-index tensors, it turns out that it is sufficient to start with

$$
\begin{aligned}
& T_{11}(1234)=\left(u_{1} \gamma^{m} u_{2}\right)\left(u_{3} \gamma^{n} u_{4}\right), \\
& T_{31}(1234)=\left(u_{1} \gamma^{a} \gamma^{m} \gamma^{n} u_{2}\right)\left(u_{3} \gamma_{a} u_{4}\right), \\
& T_{33}(1234)=\left(u_{1} \gamma^{a b} \gamma^{m} u_{2}\right)\left(u_{3} \gamma_{a b} \gamma^{n} u_{4}\right),
\end{aligned}
$$

and permutations of the spinor labels. It would be very tiresome to systematically apply a variety of Fierz transformations by hand and to find an independent set. Fortunately, by choosing a gamma matrix representation (such as the one listed in appendix B and reducing all expressions to polynomials in the independent spinor components $u_{i}^{1}, \ldots, u_{i}^{16}$, this problem can be solved with computer help. As expected, one finds that the $T_{i j}(a b c d)$ span a nine-dimensional space, and a basis can be chosen as

$$
\begin{array}{r}
T_{11}(1234), T_{11}(1324), T_{11}(1423), T_{11}(3412), T_{11}(2413), T_{11}(2314), \\
T_{31}(1234), T_{31}(1324), T_{31}(2314) . \tag{A.1}
\end{array}
$$

A typical relation reducing the other $T_{i j}(a b c d)$ to this basis is

$$
\begin{equation*}
T_{31}(3412)=2 T_{11}(1234)-2 T_{11}(3412)+T_{31}(1324)+T_{31}(3412)+2 \eta_{m n} T_{1}(1234) . \tag{A.2}
\end{equation*}
$$

Having solved the first step, it is now easy to include the two or four momenta, taking the Dirac equation into account. Consider first the case of two momenta. Starting from the two-tensors in (A.1), one gets the three independent scalars

$$
\left(u_{1} k_{3} u_{2}\right)\left(u_{3} k_{1} u_{4}\right), \quad\left(u_{1} k_{2} u_{3}\right)\left(u_{2} k_{1} u_{4}\right), \quad\left(u_{1} k_{2} u_{4}\right)\left(u_{2} k_{1} u_{3}\right)
$$

In addition, there are four products of the two independent scalars $T_{1}(1234)$ and $T_{1}(1324)$ with the two independent momentum invariants $s_{12}$ and $s_{13}$. By contracting (A.2) with momenta, one can show that

$$
\begin{align*}
& s_{12} T_{1}(1324)-s_{13} T_{1}(1234) \\
& \quad=-\left(u_{1} \not k_{3} u_{2}\right)\left(u_{3} \not k_{1} u_{4}\right)+\left(u_{1} k_{2} u_{3}\right)\left(u_{2} \not k_{1} u_{4}\right)-\left(u_{1} \not k_{2} u_{4}\right)\left(u_{2} \not k_{1} u_{3}\right) \tag{A.3}
\end{align*}
$$

and this relation can be used to eliminate $s_{12} T_{1}$ (1324). (It will become clear later that there are no independent other relations like this one.) There are thus six independent $k^{2} u_{1} \cdots u_{4}$ scalars:

$$
\begin{array}{ll}
\left(u_{1} k_{3} u_{2}\right)\left(u_{3} \not k_{1} u_{4}\right) & s_{12} T_{1}(1234) \\
\left(u_{1} k_{2} u_{3}\right)\left(u_{2} k_{1} u_{4}\right) & s_{13} T_{1}(1234)  \tag{A.4}\\
\left(u_{1} k_{2} u_{4}\right)\left(u_{2} k_{1} u_{3}\right) & s_{13} T_{1}(1324)
\end{array}
$$

Note that there is only one completely antisymmetric combination of those, given by the right hand side of (A.3). Similarly, in the case of four momenta, one finds ten independent $k^{4} u_{1} \cdots u_{4}$ scalars:

$$
\begin{array}{ll}
B_{1}=s_{12}\left(u_{1} \not k_{3} u_{2}\right)\left(u_{3} \not k_{1} u_{4}\right) & B_{2}=s_{13}\left(u_{1} \not k_{3} u_{2}\right)\left(u_{3} \not k_{1} u_{4}\right) \\
B_{3}=s_{12}\left(u_{1} k_{2} u_{3}\right)\left(u_{2} \not k_{1} u_{4}\right) & B_{4}=s_{13}\left(u_{1} \not k_{2} u_{3}\right)\left(u_{2} k_{1} u_{4}\right) \\
B_{5}=s_{12}\left(u_{1} \not k_{2} u_{4}\right)\left(u_{2} \not k_{1} u_{3}\right) & B_{6}=s_{13}\left(u_{1} k_{2} u_{4}\right)\left(u_{2} \not k_{1} u_{3}\right) \\
B_{7}=s_{12}^{2} T_{1}(1234) & B_{8}=s_{12} s_{13} T_{1}(1234)  \tag{A.5}\\
B_{9}=s_{13}^{2} T_{1}(1234) & B_{10}=s_{13}^{2} T_{1}(1324)
\end{array}
$$

Working in a gamma matrix representation, it is again simple to construct a computer algorithm which reduces any given $k^{2} u_{1} \cdots u_{4}$ or $k^{4} u_{1} \cdots u_{4}$ scalar into polynomials of the spinor and momentum components. The Dirac equation can then be solved by breaking up the sixteen-component spinors $u_{i}$ into eight-dimensional chiral spinors $u_{i}^{\mathrm{s}}$ and $u_{i}^{\mathrm{c}}$, as in eq. (B.1). One obtains polynomials in the momentum components $k_{i}^{a}$ and the independent spinor components $\left(u_{i}^{\mathrm{c}}\right)^{1 \ldots 8}$. However, a great disadvantage of this procedure is that it breaks manifest Lorentz invariance. For example, one encounters expressions which contain subsets of terms proportional to the square of a single momentum and are therefore equal to zero, but it is difficult to recognise this with a simple algorithm. The easiest solution is to choose several sets of particular vectors $k_{i}$ satisfying $k_{i}^{2}=0$ and $\sum_{i} k_{i}=0$ and to evaluate all expressions on these vectors. (By choosing integer arithmetic, one easily avoids issues of numerical accuracy.) Substituting these sets of momentum vectors in the bases (A.4) and (A.5) gives full rank six and ten respectively, showing they are indeed linearly independent.

Equipped with a computer algorithm for these basis decompositions, one finds that the symmetriser $\mathcal{S}_{4 \mathrm{~F}}$ of section 4.2,

$$
\mathcal{S}_{4 \mathrm{~F}}=4\left(1-\pi_{12}\right)\left(1-\pi_{23}\right)\left(1+\pi_{24}\right),
$$

acts on the $B_{1} \ldots B_{10}$ basis as

$$
\begin{aligned}
\mathcal{S}_{4 \mathrm{~F}} B_{1} & =8 B_{1}+4 B_{4}+4 B_{5}+4 B_{6} \\
& \ldots \\
\mathcal{S}_{4 \mathrm{~F}} B_{10} & =-8 B_{1}-16 B_{2}+8 B_{3}+16 B_{4}-8 B_{5}-16 B_{6}+8 B_{7}+24 B_{8}+24 B_{9}+16 B_{10}
\end{aligned}
$$

and has rank three.

## A. 3 Two bosons, two fermions

The combined methods of the last two sections can easily be extended to the mixed case of two bosons and two fermions. In the one-loop calculation of section 3.3, one encounters scalars of the form $k u_{1} u_{2} F_{3} F_{4}$. A basis of such objects is given by

$$
\begin{align*}
C_{1} & =\left(u_{1} \gamma^{a} u_{2}\right) k_{a}^{3} F_{b c}^{3} F_{b c}^{4} \\
C_{2} & =\left(u_{1} \gamma^{a} u_{2}\right) F_{a b}^{3} F_{b c}^{4} k_{c}^{1} \\
C_{3} & =\left(u_{1} \gamma^{a} u_{2}\right) F_{a b}^{4} F_{b c}^{3} k_{c}^{1}  \tag{A.6}\\
C_{4} & =\left(u_{1} \gamma^{a b c} u_{2}\right) F_{a b}^{3} F_{c d}^{4} k_{d}^{3} \\
C_{5} & =\left(u_{1} \gamma^{a b c} u_{2}\right) F_{a b}^{3} F_{c d}^{4} k_{d}^{1}
\end{align*}
$$

There are two combinations antisymmetric in [12] and symmetric in (34):

$$
-C_{1}+4 C_{2}+C_{4} \quad \text { and } \quad C_{2}+C_{3}
$$

Finally, there are ten independent scalars of the form $k^{3} u_{1} u_{2} F_{3} F_{4}$ (relevant to the two-loop calculation of section 4.3), and they can all be obtained by multiplication of $C_{1} \ldots C_{5}$ with the two momentum invariants $s_{12}$ and $s_{13}$.

## B. A gamma matrix representation

A convenient representation of the $\mathrm{SO}(1,9)$ gamma matrices is given by the $32 \times 32$ matrices

$$
\Gamma^{a}=\left(\begin{array}{cc}
0 & \left(\gamma^{a}\right)^{\alpha \beta} \\
\left(\gamma^{a}\right)_{\alpha \beta} & 0
\end{array}\right)
$$

where

$$
\begin{aligned}
& \left(\gamma^{0}\right)^{\alpha \beta}=\mathbf{1}_{16}=\left(\gamma^{0}\right)_{\alpha \beta}, \\
& \left(\gamma^{9}\right)^{\alpha \beta}=\left(\begin{array}{cc}
-\mathbf{1}_{8} & 0 \\
0 & \mathbf{1}_{8}
\end{array}\right)=-\left(\gamma^{9}\right)_{\alpha \beta}
\end{aligned}
$$

and $\left(\gamma^{a}\right)^{\alpha \beta}=-\left(\gamma^{a}\right)_{\alpha \beta}, a=1 \ldots 8$, is a real, symmetric $16 \times 16$ representation for the $\mathrm{SO}(8)$ Clifford algebra,

$$
\left(\gamma^{a}\right)^{\alpha \beta}=\left(\begin{array}{cc}
0 & \sigma^{a} \\
\left(\sigma^{a}\right)^{T} & 0
\end{array}\right)^{\alpha \beta}, \quad a=1 \ldots 8
$$

as given in appendix $5 . \mathrm{B}$ of 20 . The matrices $\Gamma^{a}$ satisfy the $\mathrm{SO}(1,9)$ Clifford algebra relations,

$$
\left\{\Gamma^{a}, \Gamma^{b}\right\}=2 \eta^{a b} \mathbf{1}_{32}, \quad \eta^{a b}=(+--\cdots-)
$$

and bilinears of chiral spinors (with, say, positive chirality) are constructed as

$$
\left(u \Gamma^{\left[a_{1} \ldots a_{k}\right]} v\right)=\left(u \gamma^{\left[a_{1} \ldots a_{k}\right]} v\right)=u^{\alpha}\left(\gamma^{\left[a_{1}\right.}\right)_{\alpha \beta}\left(\gamma^{a_{2}}\right)^{\beta \gamma} \ldots\left(\gamma^{\left.a_{k}\right]}\right)_{\gamma \delta} v^{\delta}
$$

This representation is particularly suitable for the calculations outlined in appendix A because it allows a simple decomposition of $\operatorname{SO}(1,9)$ spinors into $\mathrm{SO}(8)$ spinors due to its block structure:

$$
\Gamma^{0} \cdots \Gamma^{9}=\left(\begin{array}{cc}
\mathbf{1}_{16} & 0 \\
0 & -\mathbf{1}_{16}
\end{array}\right), \quad \Gamma^{1} \cdots \Gamma^{8}=\left(\begin{array}{cccc}
\mathbf{1}_{8} & 0 & 0 & 0 \\
0 & -\mathbf{1}_{8} & 0 & 0 \\
0 & 0 & \mathbf{1}_{8} & 0 \\
0 & 0 & 0 & -\mathbf{1}_{8}
\end{array}\right)
$$

Therefore, the Dirac equation for a chiral 16-component spinor $u$,

$$
\left(\gamma^{a}\right)_{\alpha \beta} \partial_{a} u^{\alpha}=0,
$$

can be solved by splitting $u$ into two chiral eight-component spinors of $\mathrm{SO}(8)$,

$$
u=\binom{u^{\mathrm{s}}}{u^{\mathrm{c}}} \quad \text { with } \quad \gamma^{1 \ldots 8}\binom{u^{\mathrm{s}}}{u^{\mathrm{c}}}=\binom{+u^{\mathrm{s}}}{-u^{\mathrm{c}}}
$$

One obtains the coupled equations

$$
\begin{aligned}
\left(\partial_{0}+\partial_{9}\right) u^{\mathrm{s}}-(\sigma \cdot \partial) u^{\mathrm{c}} & =0 \\
\left(\partial_{0}-\partial_{9}\right) u^{\mathrm{c}}-\left(\sigma^{T} \cdot \partial\right) u^{\mathrm{s}} & =0
\end{aligned}
$$

(with eight-dimensional dot products). These can be solved for $u^{\mathrm{s}}$ in terms of $u^{\mathrm{c}}$ :

$$
\begin{equation*}
u^{\mathrm{s}}=\frac{-i}{\sqrt{2} k_{+}}(\sigma \cdot \partial) u^{\mathrm{c}}=\frac{1}{\sqrt{2} k_{+}}(\sigma \cdot k) u^{\mathrm{c}} \tag{B.1}
\end{equation*}
$$

where $k_{+}=-i \partial_{+}=\frac{-i}{\sqrt{2}}\left(\partial_{0}+\partial_{9}\right)$.

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